# Axial vector transition form factors of $N \rightarrow \Delta$ in QCD 

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#### Abstract

The isovector axial vector form factors of $N \rightarrow \Delta$ transition are calculated by employing Light-cone QCD sum rules. The analytical results are analyzed by both the conventional method, and also by a Monte Carlo based approach which allows one to scan all of the parameter space. The predictions are also compared with the results in the literature, where available. Although the Monte Carlo analysis predicts large uncertainties in the predicted results, the predictions obtained by the conventional analysis are in good agreement with other results in the literature. © 2016 The Author(s). Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/4.0/). Funded by SCOAP ${ }^{3}$.


## 1. Introduction

Form factors are important properties of hadrons that contain information about their dynamics and internal structure. In this work, isovector axial form factors for the $N \rightarrow \Delta$ transition are studied. These form factors are also important in the pion production off the nucleon, where the $\Delta$ baryon appears as an intermediate resonance, and can strongly influence the production rates (see e.g. [1]). The N to $\Delta$ transition has the advantage that the $\Delta$ is the dominant, clearly accessible nucleon resonance [2]. The information contained in the axial isovector N to $\Delta$ transition

[^0]form factors can be considered as supplementary to that from electromagnetic form factors. It is also possible to check the validity of the off-diagonal Goldberger-Treiman relation using a measurement of the isovector axial vector form factors for N to $\Delta$ transition [3,4].

The first experimental results of $N \rightarrow \Delta$ axial vector form factors come from $G^{0}$ collaborations at JLAB [5] based on parity-violating electron scattering. At $Q^{2}=0.34 \mathrm{GeV}^{2}$, the obtained value of axial form factors determined from the hydrogen asymmetry in inelastic electron-proton scattering is $G_{A}^{N \Delta}=-0.05 \pm(0.35)_{\text {stat }} \pm(0.34)_{\text {sys }} \pm(0.06)_{t h}$. Furthermore, several theoretical investigations of $N \rightarrow \Delta$ axial vector form factors are carried out, on lattice QCD [2], in quark models [6], using light-cone QCD sum rules [3], chiral perturbation theory ( $\chi \mathrm{PT}$ ) [7,8] or weak single pion production [4,1,9].

To study form factors, a non-perturbative method is necessary. The light cone QCD sum rules (LCSR) is one of these non-perturbative methods that has been successfully applied to study non-perturbative hadronic properties [10-12]. In LCSR, the properties of the hadron under study are expressed in terms of the properties of the vacuum and the light cone distribution amplitudes (DAs) of the hadron.

Since the form factors are expressed in terms of the properties of the vacuum and DAs of the hadron, any uncertainty in these parameters should be reflected in the obtained predictions, on top of any other uncertainties inherent in the sum rules approach. A Monte Carlo based approach is proposed to estimate the errors due to the uncertainties of these parameters by scanning the parameter space in [13]. In this work, LCSR will be used to study the $N \rightarrow \Delta$ transition and the Monte Carlo based method will be applied to estimate the errors in the form factors by extending the method proposed in [13] to the study of form factors.

The paper is organized as follows. In section 2, the calculation of the sum rules in LCSR will be presented. In section 3, Monte Carlo analysis to estimate the uncertainties in the predictions due to the input parameters will be presented. In the last section, we conclude our work with a discussion of our results.

## 2. Formulation of baryon axial form factors

The LCSR for axial form factors of $N \rightarrow \Delta$ transition is derived from the following vacuum to nucleon correlation function:

$$
\begin{equation*}
\Pi_{\mu \nu}(p, q)=i \int d^{4} x e^{i q x}\langle 0| T\left[J_{\mu}^{\Delta}(0) A_{\nu}^{3}(x)\right]|N(p, s)\rangle, \tag{1}
\end{equation*}
$$

where $J_{\mu}^{\Delta}(x)$ is an interpolating current for the $\Delta$, and $A_{v}^{3}$ is an axial vector-isovector current defined as

$$
\begin{equation*}
A_{v}^{3}(x)=\frac{1}{2}\left(\bar{u}(x) \gamma_{\nu} \gamma_{5} u(x)-\bar{d}(x) \gamma_{\nu} \gamma_{5} d(x)\right) \tag{2}
\end{equation*}
$$

The correlation function given in Eq. (1) can be calculated in terms of hadronic properties, the so called hadronic representation, if $p^{2}>0$ and $(p+q)^{2}>0$, and also in terms of QCD parameters and several distribution amplitudes (DAs) of the $N$ baryon, the so called QCD representation, in the deep Euclidean region $p^{2} \rightarrow-\infty$ and $(p+q)^{2} \rightarrow-\infty$.

The hadronic representation of the correlation function can be obtained by inserting a complete set of states with the same quantum numbers as the interpolating currents.

$$
\begin{equation*}
\Pi_{\mu \nu}(p, q)=\sum_{s^{\prime}} \frac{\langle 0| j_{\mu}^{\Delta}\left|\Delta\left(p^{\prime}, s^{\prime}\right)\right\rangle\left\langle\Delta\left(p^{\prime}, s^{\prime}\right)\right| A_{v}^{3}|N(p, s)\rangle}{M_{\Delta}^{2}-p^{\prime 2}}+\ldots \tag{3}
\end{equation*}
$$

where $M_{\Delta}$ is the mass of the $\Delta$ baryon. The $N \rightarrow \Delta$ matrix element appearing in Eq. (3) can be written in terms of four invariant form factors as follows [14-16];

$$
\begin{align*}
& \left\langle\Delta\left(p^{\prime}, s^{\prime}\right)\right| A_{\nu}^{3}|N(p, s)\rangle \\
& \quad=i \bar{v}^{\lambda}\left(p^{\prime}, s^{\prime}\right)\left[\left\{\frac{C_{3}^{N \Delta}\left(q^{2}\right)}{M_{N}} \gamma_{\mu}+\frac{C_{4}^{N \Delta}\left(q^{2}\right)}{M_{N}^{2}} p_{\mu}^{\prime}\right\}\left(g_{\lambda \nu} g_{\rho \mu}-g_{\lambda \rho} g_{\mu \nu}\right) q^{\rho}\right. \\
& \left.\quad+C_{5}^{N \Delta}\left(q^{2}\right) g_{\lambda \nu}+\frac{C_{6}^{N \Delta}\left(q^{2}\right)}{M_{N}^{2}} q_{\lambda} q_{\nu}\right] N(p, s) \tag{4}
\end{align*}
$$

where $M_{N}$ is the mass of the nucleon, and $q=p-p^{\prime}$.
The remaining matrix element, that of the interpolating current $\langle 0| J_{\mu}^{\Delta}\left|\Delta\left(p^{\prime}, s^{\prime}\right)\right\rangle$ is defined as

$$
\langle 0| J_{\mu}^{\Delta}\left|\Delta\left(p^{\prime}, s^{\prime}\right)\right\rangle=\lambda_{\Delta} v_{\mu}\left(p^{\prime}, s^{\prime}\right)
$$

where $\lambda_{\Delta}$ is overlap amplitude of $\Delta$ baryon and $v_{\mu}\left(p^{\prime}, s^{\prime}\right)$ is the $\Delta$ baryon spinor. Summation over spins of $\Delta$ baryon is defined as:

$$
\begin{equation*}
\sum_{s^{\prime}} v_{\mu}\left(p^{\prime}, s^{\prime}\right) \overline{v_{\nu}}\left(p^{\prime}, s^{\prime}\right)=-\left(\not p^{\prime}+M_{\Delta}\right)\left[g_{\mu \nu}-\frac{1}{3} \gamma_{\mu} \gamma_{\nu}-\frac{2 p_{\mu}^{\prime} p_{v}^{\prime}}{3 M_{\Delta}^{2}}+\frac{p_{\mu}^{\prime} \gamma_{\nu}-p_{\nu}^{\prime} \gamma_{\mu}}{3 M_{\Delta}}\right] \tag{5}
\end{equation*}
$$

Inserting Eq. (4) into Eq. (3) and using Eq. (5), the correlation function Eq. (1) can be expressed in terms of the form factors as follows:

$$
\begin{align*}
\Pi_{\mu \nu}(p, q) & =-i \frac{\lambda_{\Delta}}{M_{\Delta}^{2}-p^{\prime 2}}\left[C_{3}^{N \Delta}\left(q^{2}\right)\left\{1-\frac{M_{\Delta}}{M_{N}}\right\}\left(q_{\mu} \gamma_{\nu}-g_{\mu \nu} \phi\right)\right. \\
& +\left\{C_{5}^{N \Delta}\left(q^{2}\right)+C_{4}^{A}\left(q^{2}\right) \frac{p^{\prime} \cdot q}{M_{N}^{2}}\right\} g_{\mu \nu} \phi \\
& +\frac{C_{3}^{N \Delta}\left(q^{2}\right)}{M_{N}}\left(q_{\mu} \gamma_{\nu}-g_{\mu \nu} \phi\right) \notin \\
& +\left\{-\frac{2 C_{3}^{N \Delta}\left(q^{2}\right)}{M_{N}}-\frac{C_{4}^{N \Delta}\left(q^{2}\right)}{M_{N}}\left(1+\frac{M_{\Delta}}{M_{N}}\right)\right\}\left(q_{\mu} p_{\nu}^{\prime}-g_{\mu \nu} p^{\prime} \cdot q\right) \\
& +\left\{C_{5}^{N \Delta}\left(q^{2}\right)\left(M_{\Delta}+M_{N}\right)\right\} g_{\mu \nu} \\
& +\left\{C_{6}^{N \Delta}\left(q^{2}\right)\left(\frac{M_{\Delta}+M_{N}}{M_{N}^{2}}\right)\right\} q_{\mu} q_{\nu}-\left\{\frac{C_{4}^{N \Delta}\left(q^{2}\right)}{M_{N}^{2}}\right\} q_{\mu} p_{\nu}^{\prime} \phi \\
& \left.+\left\{\frac{C_{6}^{N \Delta}\left(q^{2}\right)}{M_{N}^{2}}\right\} q_{\mu} q_{\nu} \phi\right] \tag{6}
\end{align*}
$$

In this expression, only the contribution of the spin-3/2 $\Delta$ baryon is shown. In principle, the correlation function can also receive contributions from spin- $1 / 2$ particles. The overlap of the spin-1/2 particles with the $J_{\mu}^{\Delta}$ current can be written as

$$
\begin{equation*}
\left\langle 1 / 2\left(p^{\prime}\right)\right| J_{\mu}^{\Delta}|0\rangle=\left(A p_{\mu}^{\prime}+B \gamma_{\mu}\right) u\left(p^{\prime}\right) \tag{7}
\end{equation*}
$$

where $u\left(p^{\prime}\right)$ is the spinor describing the spin- $1 / 2$ particle. Hence, if in the correlation function, the $\gamma_{\mu}$ matrix is carried to the left, only the terms that are proportional to $p_{\mu}^{\prime}$ or contain a $\gamma_{\mu}$ at the far left receive contributions from the spin- $1 / 2$ particles. Following this observation, we will order the gamma matrices as $\gamma_{\mu} \gamma_{\nu} \phi \not p^{\prime}$. The $\not p^{\prime}$ matrix can be eliminated using the equation of motion. After these steps, any structure that is not proportional to $p_{\mu}^{\prime}$, or that does not contain a $\gamma_{\mu}$ receives contributions from the spin-3/2 particles only [17,18].

To obtain the expression of the correlation function in terms of the QCD parameters and the DAs, an explicit form for the interpolating current of the $\Delta$ baryon needs to be chosen. In this work, the interpolating current is chosen as follows:

$$
\begin{equation*}
J_{\mu}^{\Delta}(x)=\frac{1}{\sqrt{3}} \epsilon^{a b c}\left[2\left(u^{a T}(x) C \gamma_{\mu} d^{b}(x)\right) u^{c}(x)+\left(u^{a T}(x) C \gamma_{\mu} u^{b}(x)\right) d^{c}(x)\right] \tag{8}
\end{equation*}
$$

Here $a, b, c$ are color indices and $C$ denotes charge conjugation. After contracting one pair of the light quark operators, the correlation function becomes:

$$
\begin{align*}
\left(\Pi_{\mu \nu}\right)_{\lambda \eta}= & \frac{i}{8 \sqrt{3}} \int d^{4} x e^{i q x}\left(C \gamma_{\mu}\right)_{\alpha \beta}\left(\gamma_{\nu} \gamma_{5}\right)_{\rho \sigma}\left\{4 \epsilon^{a b c}\langle 0| q_{1}^{a}(0) q_{2}^{b}(x) q_{3}^{c}(0)|N(p, s)\rangle\right. \\
& {\left[2 \delta_{\alpha}^{\eta} \delta_{\sigma}^{\theta} \delta_{\beta}^{\phi} S(-x)_{\lambda \rho}+2 \delta_{\lambda}^{\eta} \delta_{\sigma}^{\theta} \delta_{\beta}^{\phi} S(-x)_{\alpha \rho}\right.} \\
& \left.+\delta_{\alpha}^{\eta} \delta_{\sigma}^{\theta} \delta_{\lambda}^{\phi} S(-x)_{\beta \rho}+\delta_{\beta}^{\eta} \delta_{\sigma}^{\theta} \delta_{\phi}^{\lambda} S(-x)_{\alpha \rho}\right] \\
& -4 \epsilon^{a b c}\langle 0| q_{1}^{a}(0) q_{2}{ }_{\theta}^{b}(0) q_{3}^{c}(x)|N(p, s)\rangle \\
& {\left.\left[2 \delta_{\alpha}^{\eta} \delta_{\lambda}^{\theta} \delta_{\sigma}^{\phi} S(-x)_{\beta \rho}+\delta_{\alpha}^{\eta} \delta_{\beta}^{\theta} \delta_{\sigma}^{\phi} S(-x)_{\lambda \rho}\right]\right\} } \tag{9}
\end{align*}
$$

where $\lambda$ and $\eta$ are spinor indices, and $S(x)$ represents the light-quark propagator and is given by:

$$
\begin{align*}
S_{q}(x) & =\frac{i \not x}{2 \pi^{2} x^{4}}-\frac{\langle q \bar{q}\rangle}{12}\left(1+\frac{m_{0}^{2} x^{2}}{16}\right) \\
& -i g_{s} \int_{0}^{1} d v\left[\frac{\not x}{16 \pi^{2} x^{4}} G_{\mu \nu} \sigma^{\mu \nu}-v x^{\mu} G_{\mu \nu} \gamma^{\nu} \frac{i}{4 \pi^{2} x^{2}}\right] . \tag{10}
\end{align*}
$$

In this expression, $\langle q \bar{q}\rangle$ is the quark condensate, $m_{0}$ is defined in terms of the mixed quark gluon condensate as $\left\langle\bar{q} g_{s} G^{\mu \nu} \sigma_{\mu \nu} q\right\rangle \equiv m_{0}^{2}\langle\bar{q} q\rangle$ and $g_{s}$ is the strong coupling constant. The terms proportional to $G_{\mu \nu}$ arise from the interaction of the propagating quark with the external gluonic field and lead to contributions from the four-particle nucleon distribution amplitudes. Such corrections from higher Fock space components of the nucleon wave function are not expected to play an important role [19], and hence we would not take them into account. Additionally, the terms proportional to quark condensates are removed by Borel transformations. Hence the first term, which contains the hard light-quark propagator, will be considered for our discussion. To proceed with the calculation of the correlation function, the matrix element of the local three-quark operator $4 \epsilon^{a b c}\langle 0| q_{1 \alpha}^{a}\left(a_{1} x\right) q_{2 \beta}^{b}\left(a_{2} x\right) q_{3 \gamma}^{c}\left(a_{3} x\right)|N(p, s)\rangle$ is needed. The light-cone dis-
tribution amplitudes of the nucleon, which we use in our work to extract the axial form factors, are presented in Ref. [20] up to twist six on the basis of QCD conformal partial wave expansion. The expansion is in terms of increasing twist, where the twist of a DA is defined as the difference between the dimension and spin of the operators contributing to that DA. We refer the reader to Ref. [20] for a detailed analysis of the distribution amplitudes of the nucleon. Using the most general decomposition of the matrix element (see Eq. (2.3) in Ref. [21]) and taking the Fourier transformations appearing in Eq. (9), the QCD representation of the correlation function can be obtained.

Note that the hadronic representation, Eq. (6), and the QCD representation are obtained in different kinematical regions. The two expressions can be related to each other by using the spectral representation of the correlation functions. Quite generally, the coefficients of various structures in the correlation function can be written as:

$$
\begin{equation*}
\Pi\left(p^{2}, p^{\prime 2} ; Q^{2}\right)=\int_{0}^{\infty} d s_{1} d s_{2} \frac{\rho\left(s_{1}, s_{2} ; Q^{2}\right)}{\left(s_{1}-p^{2}\right)\left(s_{2}-p^{\prime 2}\right)}+\text { polynomials in } p^{2} \text { or } p^{\prime 2} \tag{11}
\end{equation*}
$$

where $\rho$ is called the spectral density. The spectral density can be calculated both using the hadronic representation of the correlation function, $\rho^{h}$, or using the QCD representation, $\rho^{Q C D}$. Once $\rho$ is obtained, the spectral representation allows one to evaluate the correlation function in all kinematical regions for $p^{2}$ and $p^{\prime 2}$.

The LCSR are obtained by matching the expression of the correlation function in terms of QCD parameters to its expression in terms of the hadronic properties, using their spectral representation. In order to do this, we choose the structures proportional to $q_{\mu} \gamma_{\nu} q^{\prime}, q_{\mu} p^{\prime}{ }_{\nu} \phi$, $\left(g_{\mu \nu} \phi-q_{\mu} \gamma_{\nu} \phi\right)$ and $q_{\mu} q_{\nu} \phi$ for the form factors $C_{3}^{N \Delta}, C_{4}^{N \Delta}, C_{5}^{N \Delta}$ and $C_{6}^{N \Delta}$, respectively. For the $N \rightarrow \Delta$ transition form factors, we obtain:

$$
\begin{align*}
C_{3}^{N \Delta}\left(Q^{2}\right) \frac{\lambda_{\Delta}}{M_{\Delta}^{2}-p^{\prime 2}} & =-\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d \alpha \frac{(1-\alpha)}{(q-p \alpha)^{4}}\left[F_{9}(\alpha)-F_{10}(\alpha)\right] \\
& +\frac{M_{N}}{\sqrt{3}} \int_{0}^{1} d x_{i} \frac{1}{\left(q-p x_{i}\right)^{2}}\left[2 F_{11}\left(x_{i}\right)-F_{12}\left(x_{i}\right)\right] \\
& -\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d x_{i} \frac{1}{\left(q-p x_{i}\right)^{4}}\left[2 F_{13}\left(x_{i}\right)-F_{14}\left(x_{i}\right)\right] \\
& -\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d \beta \frac{1}{(q-p \beta)^{4}}\left[2 F_{15}(\beta)-F_{16}(\alpha)\right]  \tag{12}\\
C_{4}^{N \Delta}\left(Q^{2}\right) \frac{\lambda_{\Delta}}{M_{\Delta}^{2}-p^{\prime 2}} & =-\frac{M_{N}^{5}}{\sqrt{3}} \int_{0}^{1} d \beta \frac{(1-\beta)^{2}}{(q-p \beta)^{6}}\left[4 F_{5}(\beta)-2 F_{7}(\beta)\right] \\
& +\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d \alpha \frac{(1-\alpha)}{(q-p \alpha)^{4}}\left[2 F_{6}(\alpha)+F_{8}(\alpha)\right] \tag{13}
\end{align*}
$$

$$
\begin{align*}
& C_{5}^{N \Delta}\left(Q^{2}\right) \frac{\lambda_{\Delta}}{M_{\Delta}^{2}-p^{\prime 2}}=\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d \beta \frac{1}{(q-p \beta)^{4}}\left[F_{17}(\beta)-F_{18}(\beta)\right] \\
& -\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d x_{i} \frac{1}{\left(q-p x_{i}\right)^{4}}\left[F_{19}\left(x_{i}\right)+F_{20}\left(x_{i}\right)\right] \\
& +\frac{M_{N}}{\sqrt{3}} \int_{0}^{1} d x_{i} \frac{1}{\left(q-p x_{i}\right)^{2}}\left[F_{21}\left(x_{i}\right)-F_{22}\left(x_{i}\right)\right] \\
& -\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d \beta \frac{2}{(q-p \beta)^{4}}\left[F_{23}(\beta)-F_{24}(\beta)\right] \\
& +\frac{M_{N}}{\sqrt{3}} \int_{0}^{1} d x_{i} \frac{1-x_{i}}{\left(q-p x_{i}\right)^{2}}\left[F_{25}\left(x_{i}\right)-F_{26}\left(x_{i}\right)\right] \\
& -\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d x_{i} \frac{1}{\left(q-p x_{i}\right)^{4}}\left[F_{27}\left(x_{i}\right)-F_{28}\left(x_{i}\right)\right] \\
& +\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d \beta \frac{1-\beta}{(q-p \beta)^{4}}\left[F_{29}(\beta)-F_{30}(\beta)\right] \\
& +\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d \alpha \frac{1-\alpha}{(q-p \alpha)^{4}}\left[F_{31}(\alpha)-F_{32}(\alpha)\right] \\
& -\frac{M_{N}}{\sqrt{3}} \int_{0}^{1} d x_{i} \frac{1}{\left(q-p x_{i}\right)^{2}}\left[2 F_{33}\left(x_{i}\right)-F_{34}\left(x_{i}\right)\right] \\
& +\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d x_{i} \frac{1}{\left(q-p x_{i}\right)^{4}}\left[2 F_{35}\left(x_{i}\right)+F_{36}\left(x_{i}\right)\right] \\
& +\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d \beta \frac{1}{(q-p \beta)^{2}}\left[2 F_{37}(\beta)-F_{38}(\beta)\right] \\
& +\frac{M_{N}^{3}}{\sqrt{3}} \int_{0}^{1} d \beta \frac{1-\beta}{(q-p \beta)^{4}}\left[2 F_{39}(\beta)+F_{40}(\beta)\right] \\
& -\frac{M_{N}}{\sqrt{3}} \int_{0}^{1} d \alpha \frac{1}{(q-p \alpha)^{2}}\left[F_{41}(\alpha)+F_{42}(\alpha)\right] \tag{14}
\end{align*}
$$

$$
\begin{align*}
C_{6}^{N \Delta}\left(Q^{2}\right) \frac{\lambda_{\Delta}}{M_{\Delta}^{2}-p^{\prime 2}} & =\frac{M_{N}^{5}}{\sqrt{3}} \int_{0}^{1} d \beta \frac{(1-\beta)^{2}}{(q-p \beta)^{6}}\left[4 F_{1}(\beta)+2 F_{3}(\beta)\right] \\
& -\frac{M_{N}^{3}}{\sqrt{3}} \int_{o}^{1} d \alpha \frac{(1-\alpha)}{(q-p \alpha)^{4}}\left[F_{2}(\alpha)-F_{4}(\alpha)\right] \tag{15}
\end{align*}
$$

The explicit forms of the functions that appear in the above sum rules in terms of the DAs of the nucleon are given in Appendix A. In these expressions, the left hand side is actually a sum over the contributions of all spin- $3 / 2$ baryons. The sum rules are obtained by carrying out Borel transformation to eliminate any polynomials that arise during the matching. Furthermore, Borel transformation also suppresses the contributions of higher states and continuum. After Borel transformations, Eqs. (12)-(15) take the form

$$
\begin{equation*}
C_{i}^{N \Delta}\left(Q^{2}\right) e^{-\frac{m_{\Delta}^{2}}{M^{2}}}=\int_{0}^{1} d x \rho^{i}(x) e^{-\frac{s(x)}{M^{2}}} \quad i=3,4,5, \text { or } 6 \tag{16}
\end{equation*}
$$

where

$$
s(x)=(1-x) M_{N}^{2}+\frac{1-x}{x} Q^{2},
$$

with $Q^{2}=-q^{2}$ and $M$ is the Borel parameter.

## 3. Traditional analysis vs. Monte Carlo analysis

In the traditional analysis of sum rules, the spectral density of the higher states and the continuum are parameterized using quark hadron duality. In this approach, it is assumed that $\rho^{h}(s)=\rho^{Q C D}(s)$ when $s>s_{0}$, i.e. the contribution of the higher states and continuum to the spectral density is approximated by the spectral density expressed in terms of the QCD parameters. Both the Borel transformation and the subtraction of the higher states and the continuum are carried out using the following substitution rules (see e.g. [20]):

$$
\begin{align*}
\int d x \frac{\rho(x)}{(q-x p)^{2}} \rightarrow & -\int_{x_{0}}^{1} \frac{d x}{x} \rho(x) e^{-s(x) / M^{2}}, \\
\int d x \frac{\rho(x)}{(q-x p)^{4}} \rightarrow & \frac{1}{M^{2}} \int_{x_{0}}^{1} \frac{d x}{x^{2}} \rho(x) e^{-s(x) / M^{2}}+\frac{\rho\left(x_{0}\right)}{Q^{2}+x_{0}^{2} M_{N}^{2}} e^{-s_{0} / M^{2}}, \\
\int d x \frac{\rho(x)}{(q-x p)^{6}} \rightarrow & -\frac{1}{2 M^{4}} \int_{x_{0}}^{1} \frac{d x}{x^{3}} \rho(x) e^{-s(x) / M^{2}}-\frac{1}{2 M^{2}} \frac{\rho\left(x_{0}\right)}{x_{0}\left(Q^{2}+x_{0}^{2} M_{N}^{2}\right)} e^{-s_{0} / M^{2}} \\
& +\frac{1}{2} \frac{x_{0}^{2}}{Q^{2}+x_{0}^{2} M_{N}^{2}}\left[\frac{d}{d x_{0}} \frac{\rho\left(x_{0}\right)}{x_{0}\left(Q^{2}+x_{0}^{2} M_{N}^{2}\right)}\right] e^{-s_{0} / M^{2}}, \tag{17}
\end{align*}
$$

where $x_{0}$ is the solution of the quadratic equation for $s=s_{0}$ :

$$
x_{0}=\left[\sqrt{\left(Q^{2}+s_{0}-M_{N}^{2}\right)^{2}+4 M_{N}^{2} Q^{2}}-\left(Q^{2}+s_{0}-M_{N}^{2}\right)\right] /\left(2 M_{N}^{2}\right)
$$

where $s_{0}$ is the continuum threshold. $s_{0} \rightarrow \infty\left(x_{0}=0\right)$ limit corresponds to Borel transformation without subtraction.

An alternative approach to analyze the sum rules using Monte Carlo methods is presented in [13]. The method has the advantage that it allows for a more reliable estimate of the error bars. In this method, one chooses random values for the uncertain parameters appearing in the sum rules within their error bars. For each set of values of the parameters, one obtains a numerical prediction for quantity under study. The obtained results are then analyzed statistically to obtain the central value and the error bars (see [13] for details). In [13], the method has been used for mass sum rules, and in [22,23], it has been applied to calculating coupling constants. Here we generalize that method to analyze the form factors. The steps of the analysis can be summarized as follows:

1. Let $p_{i}$ denote one of the input parameters whose central value is $\bar{p}_{i}$ and whose error is $\sigma_{i}$. We will assume that $p_{i}$ is a normally distributed random variable with variance $\sigma_{i}$. Let $\left\{p_{k}\right\}$ denote a possible set of input parameters. Choose $N$ such sets. In this work, $N$ is chosen to be $N=1000$.
2. Let $\Pi_{\left\{p_{k}\right\}}\left(Q^{2}, M^{2}\right)$ be the correlation function calculated using the set $\left\{p_{k}\right\}$ of input parameters for the fixed values of $M^{2}$ and $Q^{2}$. Denote the average and the standard deviation of these values by $\bar{\Pi}\left(Q^{2}, M^{2}\right)$ and $\sigma_{\Pi}\left(Q^{2}, M^{2}\right)$.
3. For each set $\left\{p_{k}\right\}$ of parameters, define $\chi_{\left\{p_{k}\right\}}^{2}$ distribution for a fixed $Q^{2}$ as

$$
\begin{equation*}
\chi_{\left\{p_{k}\right\}}^{2}\left(Q^{2}\right)=\sum_{M^{2}} \frac{\left(\Pi_{\left\{p_{k}\right\}}\left(Q^{2}, M^{2}\right)-\Pi^{\text {model }}\left(Q^{2}, M^{2} ; C\left(Q^{2}\right), a_{n}\right)\right)^{2}}{\sigma_{\Pi}\left(Q^{2}, M^{2}\right)^{2}} \tag{18}
\end{equation*}
$$

where the model for the correlation function depends on the form factor $C\left(Q^{2}\right)$ and possibly other parameters $a_{n}$.
4. Minimizing $\chi_{\left\{p_{k}\right\}}^{2}\left(Q^{2}\right)$ with respect to the form factor $C\left(Q^{2}\right)$ and other parameter of the model for the correlation function yields the form factor for the given set of parameter $C_{\left\{p_{k}\right\}}\left(Q^{2}\right)$.
5. Repeating this procedure for all the parameter sets, a distribution for the value of the form factor at the chosen value of $Q^{2}$ is obtained. The prediction for the value of the form factor at the chosen value of $Q^{2}$ can be taken as the average of this distribution and the uncertainty as the standard deviation of this distribution.
6. This procedure is repeated for various values of $Q^{2}$ to obtain the value and the uncertainty of the value of the form factor can be obtained.

In this work, the correlation has been modelled as

$$
\begin{equation*}
\Pi=a_{0} C^{N \Delta}\left(Q^{2}\right) e^{-\frac{M_{\Delta}^{2}}{M^{2}}}+a_{1} e^{-\frac{m_{1}^{2}}{M^{2}}}+a_{2} e^{-\frac{m_{2}^{2}}{M^{2}}} \tag{19}
\end{equation*}
$$

where we imposed the constraint $m_{2}>m_{1}>M_{\Delta}$. Both a triple exponential (TE) and a double exponential (DE) fit (by setting $a_{2}=0$ ) have been performed.

Table 1
The values of the parameters entering the DAs of $N$. The upper panel shows the dimensionful parameters N . In the lower panel we list the values of the five parameters that determine the shape of the DAs.

| $f_{N}\left(\mathrm{GeV}^{2}\right)$ | $\lambda_{1}\left(\mathrm{GeV}^{2}\right)$ | $\lambda_{2}\left(\mathrm{GeV}^{2}\right)$ |
| :--- | :--- | :--- |
| $0.005 \pm 0.0005$ | $-0.027 \pm 0.009$ | $0.054 \pm 0.019$ |


| $V_{1}^{d}$ | $A_{1}^{u}$ | $f_{1}^{d}$ | $f_{2}^{d}$ | $f_{1}^{u}$ |
| :--- | :--- | :--- | :--- | :--- |
| $0.23 \pm 0.03$ | $0.38 \pm 0.15$ | $0.40 \pm 0.05$ | $0.22 \pm 0.05$ | $0.07 \pm 0.05$ |

## 4. Results and conclusion

To obtain a numerical prediction for the form factors, first one needs the expressions for the distribution amplitudes of the nucleon [20], the distribution amplitudes can be expressed in terms of several parameters. Not all of the parameters are independent. The independent parameters and their numerical values are presented in Table 1 [20]. Another non-perturbative parameter that is required is the residue of the $\Delta$ baryon $\lambda_{\Delta}$. This residue is obtained from the mass sum rules for the $\Delta$ baryons to be $\lambda_{\Delta}=0.038 \mathrm{GeV}^{3}$ [24-26].

In the conventional analysis, the obtained predictions for the form factors depend on two auxiliary parameters: the Borel parameter $M^{2}$, and the continuum threshold $s_{0}$. The continuum threshold signals the energy the scale at which, the excited states and continuum start to contribute to the correlation function. There are various proposal on how to determine this parameter. One approach is to vary this parameter in a reasonable range, until a Borel window appears in which the predictions are independent of the Borel parameter [27]. Another recent proposal is to choose $s_{0}$ as a function of the Borel parameter and to determine the functional dependence by requiring the independence of the mass prediction on the Borel parameter [28]. The mass of the $\Delta$ baryon can be obtained from Eq. (16) using the relation:

$$
\begin{equation*}
M_{\Delta}^{2}=M^{4} \frac{\partial}{\partial M^{2}} \ln \left(C_{i}^{N \Delta}\left(Q^{2}\right) e^{-\frac{M_{\Delta}^{2}}{M^{2}}}\right) \tag{20}
\end{equation*}
$$

But the generally accepted rule of thumb to determine this parameter is to assume that $s_{0} \simeq$ $\left(M_{\Delta}+0.3 \mathrm{GeV}\right)^{2}$ and check the dependence of the results on slight variations of this parameter. To estimate the reliability of this determination of $s_{0}$, in Fig. 1, we present the $M_{\Delta}$ predictions obtained from the form factor sum rules for $s_{0}=2.0 \mathrm{GeV}^{2}, s_{0}=2.5 \mathrm{GeV}^{2}$ and $s_{0}=3.0 \mathrm{GeV}^{2}$ at $Q^{2}=2.0 \mathrm{GeV}^{2}$. As can be seen from the figures, the predictions for the mass of $M_{\Delta}$ are within about $20 \%$ for all of the sum rules. Also, as can also be seen from the figures, within the considered range of $M^{2}$, the mass predictions are almost independent of the Borel parameter. Hence, a possible $M^{2}$ dependence of $s_{0}$ is not necessary within this range.

The Borel parameter is an unphysical parameter, and the predictions on the form factors should be independent of the value of this parameter. Due to the approximations made to obtain the sum rules, a residual dependence on $M^{2}$ remains and for this reason, a region in which the predictions are practically independent of the value of the Borel parameter needs to be chosen. In Fig. 2, we plot the dependencies of the form factors on $M^{2}$ for two fixed values of $Q^{2}$ and for various values of $s_{0}$ in the range $2.0 \mathrm{GeV}^{2} \leq s_{0} \leq 4 \mathrm{GeV}^{2}$. As can be seen from these figures, for $M^{2} \geq 3.0 \mathrm{GeV}^{2}$, there is negligible dependence of the Borel parameter. Hence, in the traditional analysis, this range of $M^{2}$ is used.


Fig. 1. The $M_{\Delta}$ predictions obtained from the form factor sum rules for $s_{0}=2.0 \mathrm{GeV}^{2}, s_{0}=2.5 \mathrm{GeV}^{2}$ and $s_{0}=$ $3.0 \mathrm{GeV}^{2}$ at $Q^{2}=2.0 \mathrm{GeV}^{2}$, (a) for $C_{3}^{N \Delta}$ form factors, (b) for $C_{4}^{N \Delta}$ form factors, (c) for $C_{5}^{N \Delta}$ form factors, (d) for $C_{6}^{N \Delta}$ form factors.

To analyze the convergence of the twist expansion, in Fig. 3, the contribution from definite twists to the form factors is presented at fixed $Q^{2}=2.0 \mathrm{GeV}^{2}$ and $s_{0}=2.5 \mathrm{GeV}^{2}$ values. As can be seen from the figures, the dominant contribution to the form factors is from twist-4 DAs. Form factors $C_{4}^{N \Delta}, C_{5}^{N \Delta}$ and $C_{6}^{N \Delta}$ receive negligible contributions from other DAs. In the case of the $C_{3}^{N \Delta}$ form factors, the contributions of all the twists are comparable, and hence, one can not talk about the convergence of the twist expansion. Note that, for the case of $C_{3}^{N \Delta}$, although the twist expansion does not seem to converge, the mass predicted by the sum rules for the $C_{3}^{N \Delta}$ form factor is quite accurate as can be seen from Fig. 1(a). Indeed, the prediction of $M_{\Delta}$ using the sum rules for $C_{3}^{N \Delta}$ is more accurate than the predictions obtained from the sum rules for $C_{4}^{N \Delta}$ and $C_{5}^{N \Delta}$. This can be interpreted as an indication in favor of the reliability of the predictions on the form factors for $C_{3}^{N \Delta}$. The baryon mass corrections to the DAs contribute only to the form factors $C_{3}^{N \Delta}$ and $C_{5}^{N \Delta}$. The contribution of these corrections to $C_{5}^{N \Delta}$ is negligible, whereas, the contribution to $C_{3}^{N \Delta}$ is twice the final result. This means that the baryon mass correction contribution to this form factors even changes the sign of the form factor and hence can not be neglected.

In Fig. 4, we present the $Q^{2}$ dependence of the form factors obtained using two different types of analysis. The results of the conventional sum rules analysis are presented with lines for the


Fig. 2. The dependence of the form factors; on the Borel parameter squared $M^{2}$ for the values of the continuum threshold $s_{0}=2.0 \mathrm{GeV}^{2}, s_{0}=2.5 \mathrm{GeV}^{2}, s_{0}=3.0 \mathrm{GeV}^{2}, s_{0}=3.5 \mathrm{GeV}^{2}$, $s_{0}=4.0 \mathrm{GeV}^{2}$ and $Q^{2}=2.0 \mathrm{GeV}^{2}, 4.0 \mathrm{GeV}^{2}$ (a) and (b) for $C_{3}^{N \Delta}$ form factor, (c) and (d) for $C_{4}^{N \Delta}$ form factor, (e) and (f) for $C_{5}^{N \Delta}$ form factor and (g) and (h) for $C_{6}^{N \Delta}$ form factor. (Note the different scales used for the vertical axis.)


Fig. 3. The contribution to the form factors at $Q^{2}=2.0 \mathrm{GeV}^{2}$ from various twists. (a) for $C_{3}^{N \Delta}$ form factors, (b) for $C_{4}^{N \Delta}$ form factors, (c) for $C_{5}^{N \Delta}$ form factors, (d) for $C_{6}^{N \Delta}$ form factors. The T3, T4, T5, T6 and M are twist-3, twist-4, twist-5, twist-6 and mass correction contributions, respectively.
central values of the parameters appearing in Table 1. The circles and squares with error bars are the results of the Monte Carlo analysis. The results obtained by both a double exponential (DE) and a triple exponential (TE) model of the correlation function are presented. It is observed that for the form factors $C_{3}^{N \Delta}\left(Q^{2}\right), C_{4}^{N \Delta}\left(Q^{2}\right)$ and $C_{6}^{N \Delta}\left(Q^{2}\right)$, Monte Carlo analysis and the prediction for the conventional sum rules analysis agree with each other at large values of $Q^{2}$, but deviate from each other for small values of $Q^{2}$. In the case of $C_{5}^{N \Delta}\left(Q^{2}\right)$, we observe that predictions for the central values agree at small $Q^{2}$, but deviate from each other for large values of $Q^{2}$. Also, it is seen that although with the conventional sum rules analysis, one predicts value for form factor $C_{3}^{N \Delta}\left(Q^{2}\right)$ that is very close to zero, the Monte Carlo analysis shows that this form factor is consistent with significant non-zero values. In the case of $C_{4}^{N \Delta}\left(Q^{2}\right)$, although the conventional sum rules analysis leads to a value that is significantly away from zero, the Monte Carlo analysis shows that this form factor is almost consistent with zero.

The values of the form factors at zero momentum transfer $Q^{2}=0$ define the corresponding charges. The sum rules method is only reliable at large enough values of $Q^{2}$, which is typically assumed to be $Q^{2}>1 \mathrm{GeV}^{2}$. To obtain the value at $Q^{2}=0$, the predictions have to be extrapolated.

The function that is typically used in the literature for the extrapolation has the form (see e.g. [4])

$$
\begin{equation*}
C_{i}^{N \Delta}\left(Q^{2}\right)=\frac{C_{i}^{N \Delta}(0)}{\left(1+Q^{2} / m_{A}^{2}\right)^{n}} \tag{21}
\end{equation*}
$$



Fig. 4. The dependence of the form factors on $Q^{2}$. The results of the Monte Carlo analysis are shown by symbols and error bars. Results of conventional analysis (calculated at $M^{2}=3.0 \mathrm{GeV}^{2}$ ) are shown by lines. (a) for $C_{3}^{N \Delta}$ form factor, (b) for $C_{4}^{N \Delta}$ form factor, (c) for $C_{5}^{N \Delta}$ form factor and (d) for $C_{6}^{N \Delta}$ form factor.
where $n=2$. The axial mass $m_{A}$ is the free parameter which has been found experimentally to be [29]

$$
\begin{equation*}
m_{A}=1.28_{-0.10}^{+0.08} \tag{22}
\end{equation*}
$$

This value is used to fit our predictions to obtain $C_{i}^{N \Delta}, i=3,4,5$ or 6 . The obtained predictions on the axial charges are presented in Table 2.

Although this function gives a reasonable fit to our predictions on $C_{3}^{N \Delta}, C_{4}^{N \Delta}$ and $C_{5}^{N \Delta}$ using the conventional analysis, it fits poorly to the form factor $C_{6}^{N \Delta}$. For this form factor, a better fit is obtained for $n=4$. Furthermore, close to $Q^{2}=1 \mathrm{GeV}^{2}$, the predictions of the form factor from sum rules deviate from a pole form significantly. This can be interpreted as a failure of the sum rules at such low values of $Q^{2}$. For this reason, we fit the form factors only in the region $Q^{2}>2 \mathrm{GeV}^{2}$.

At $Q^{2}=0$, experimentally the most easily accessible form factor is $C_{5}^{N \Delta}$ [4]. This form factor also determines the axial charge of the transition $N-\Delta$. Various predictions for this axial charge are as follows: the off-diagonal Goldberger-Treiman relation predicts that the axial charge is $C_{5}^{N \Delta}(0) \simeq 1.20$, the prediction of quark model ranges from $C_{5}^{N \Delta}(0)=0.81$ to 1.53 (see e.g. [6] and reference therein), in the case of chiral perturbation theory it is estimated to be

Table 2
The values of dipole fit parameters, $C_{i}^{N \Delta}(0)$ for axial form factor obtained from the conventional analysis, MC analysis using DE (MCDE) and MC analysis using TE (MCTE).

| Form factor | $C_{i}^{N \Delta}(0)\left(\mathrm{GeV}^{-2}\right)$ <br> (conventional analysis) | $C_{i}^{N \Delta}(0)\left(\mathrm{GeV}^{-2}\right)$ <br> $(\mathrm{MCDE})$ | $C_{i}^{N \Delta}(0)\left(\mathrm{GeV}^{-2}\right)$ <br> $(\mathrm{MCTE})$ |
| :--- | :--- | :--- | :--- |
| $C_{3}$ | $0.11 \pm 0.03$ | $0.41 \pm 0.12$ | $0.22 \pm 0.07$ |
| $C_{4}$ | $0.27 \pm 0.09$ | $0.13 \pm 0.02$ | $0.10 \pm 0.02$ |
| $C_{5}$ | $1.14 \pm 0.20$ | $0.59 \pm 0.11$ | $0.45 \pm 0.09$ |
| $C_{6}$ | $-1.65 \pm 0.46$ | $-1.88 \pm 0.15$ | $-1.96 \pm 0.16$ |

$C_{5}^{N \Delta}(0)=1.16$ [7], lattice QCD predicts $C_{5}^{N \Delta}(0)=0.9 \pm 0.02$ [2] and the results from weak pion production are $C_{5}^{N \Delta}(0)=1.08 \pm 0.1$ [4] and $C_{5}^{N \Delta}(0)=1.19 \pm 0.08$ [9]. Although these values are consistent with our prediction on $C_{5}^{N \Delta}(0)$ using the conventional analysis within error bars, the Monte Carlo analysis predicts smaller values. We can also compare our results with Adler's model predictions [14]. There, the axial form factors for the $\mathrm{N} \rightarrow \Delta$ transition have been parameterized as

$$
\begin{equation*}
C_{j}^{A}\left(Q^{2}\right)=\frac{C_{j}^{A}(0)\left(1-a_{j} Q^{2} /\left(b_{f}-Q^{2}\right)\right)}{\left(1-Q^{2} / m_{A}^{2}\right)^{2}} ; \quad j=3,4,5 \tag{23}
\end{equation*}
$$

with

$$
\begin{aligned}
& C_{3}^{N \Delta}(0)=0, C_{4}^{N \Delta}(0)=-\frac{C_{5}^{N \Delta}(0)}{4}=-0.3 \text { and } C_{5}^{N \Delta}(0)=1.2 \\
& a_{3}=b_{3}=0, a_{4}=a_{5}=-1.21, b_{4}=b_{5}=2 \mathrm{GeV}^{2} .
\end{aligned}
$$

We see the predictions of the conventional analysis are consistent with the predictions of Adler's model except of the $C_{3}^{N \Delta}$, however, the predictions of the Monte Carlo methods are not consistent.

In [5], $G_{A}^{N \Delta}$, defined as:

$$
\begin{equation*}
G_{A}^{N \Delta}\left(Q^{2}\right)=\frac{1}{2}\left[M_{N}^{2}-M_{\Delta}^{2}+Q^{2}\right] C_{4}^{A}\left(Q^{2}\right)-M_{N}^{2} C_{5}^{A}\left(Q^{2}\right), \tag{24}
\end{equation*}
$$

is measured to be $G_{A}^{N \Delta}\left(Q^{2}=0.34 \mathrm{GeV}^{2}\right)=-0.05 \pm(0.35)_{\text {stat }} \pm(0.34)_{\text {sys }} \pm(0.06)_{\text {theory }}$. Using the values in Table 2, $G_{A}^{N \Delta}=-1.04 \pm 0.19, G_{A}^{N \Delta}=-0.54 \pm 0.10, G_{A}^{N \Delta}=-0.41 \pm 0.08$ respectively using conventional analysis, MC analysis with DE, and MC analysis with TE. Note that, the experimental measurement has large error bars, but is consistent with $G_{A}^{N \Delta}=0$. The predictions of both the conventional analysis and MC analysis are consistent with the experimental value of $G_{A}^{N \Delta}$ (within error bars), but the MC analysis predicts a much smaller value for $G_{A}^{N \Delta}$.

In conclusion, we have extracted the isovector axial-vector form factors of $\mathrm{N}-\Delta$ transition by applying the LCSR. The $Q^{2}$ dependence of form factors is obtained using the conventional analysis, and Monte Carlo analysis. The Monte Carlo analysis showed that the error bars in the predictions of the form factors are large especially in the low $Q^{2}$ region. This especially implies that the extrapolation to $Q^{2}=0$ is unreliable. On the other hand, it is shown that the predictions obtained by the conventional analysis are consistent with the results in the literature, when they exist.

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## Appendix A. Explicit forms of the functions $F_{i}$ for the $N \rightarrow \Delta$ transition

$$
\begin{aligned}
& F_{1}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{2} \int_{0}^{1-x_{2}} d x_{1} 2\left[T_{1}-T_{3}-T_{4}+T_{6}-T_{7}-T_{8}-A_{1}+A_{2}-A_{3}-A_{4}\right. \\
& \left.+A_{5}-A_{6}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{2}=\int_{\alpha}^{1} d x_{2} \int_{o}^{1-x_{2}} d x_{1} 4\left[A_{1}-A_{2}+A_{3}-T_{1}+T_{3}+T_{7}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{3}=\int_{o}^{\beta} d \alpha \int_{\alpha}^{1} d x_{3} \int_{o}^{1-x_{3}} d x_{1} 2\left[V_{1}-V_{2}-V_{3}-V_{4}-V_{5}+V_{6}+T_{1}-T_{3}-T_{4}+T_{6}\right. \\
& \left.-T_{7}-T_{8}+A_{1}-A_{2}+A_{3}+A_{4}-A_{5}+A_{6}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
& F_{4}=\int_{\alpha}^{1} d x_{3} \int_{0}^{1-x_{3}} d x_{1} 2\left[V_{1}-V_{2}-V_{3}+A_{1}-A_{2}+A_{3}+T_{1}-T_{3}-T_{7}\right] \\
& \left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right) \text {, } \\
& F_{5}=\int_{o}^{\beta} d \alpha \int_{\alpha}^{1} d x_{3} \int_{o}^{1-x_{3}} d x_{1} 2\left[T_{1}-T_{3}-T_{4}+T_{6}-T_{7}-T_{8}-A_{1}+A_{2}-A_{3}-A_{4}\right. \\
& \left.+A_{5}-A_{6}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
& F_{6}=\int_{\alpha}^{1} d x_{3} \int_{\alpha}^{1-x_{3}} d x_{1} 2\left[T_{1}-T_{3}-T_{7}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
& F_{7}=\int_{o}^{\beta} d \alpha \int_{\alpha}^{1} d x_{2} \int_{o}^{1-x_{2}} d x_{1} 2\left[-A_{1}+A_{2}-A_{3}-A_{4}+A_{5}-A_{6}+T_{1}-T_{3}-T_{4}+T_{6}\right. \\
& \left.-T_{7}-T_{8}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{8}=\int_{\alpha}^{1} d x_{2} \int_{\alpha}^{1-x_{2}} d x_{1}\left[2 V_{1}-2 V_{2}-2 V_{3}+2 A_{1}-2 A_{2}+2 A_{3}-4 T_{1}+4 T_{3}+4 T_{7}\right] \\
& \left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right) \text {, } \\
& F_{9}=\int_{\alpha}^{1} d x_{2} \int_{0}^{1-x_{2}} d x_{1}\left[2 T_{1}-T_{3}-T_{4}-T_{7}-T_{8}+A_{3}-A_{4}-S_{1}+S_{2}+P_{1}-P_{2}-2 V_{1}\right.
\end{aligned}
$$

$$
\begin{aligned}
& \left.+2 V_{2}+V_{3}+V_{4}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
F_{10}= & \int_{\alpha}^{1} d x_{3} \int_{0}^{1-x_{3}} d x_{1}\left[-V_{1}+V_{3}+V_{5}+T_{1}-T_{3}-T_{7}-1 / 2 A_{1}-1 / 2 A_{3}-1 / 2 A_{5}\right] \\
& \left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
F_{11}= & \int_{0}^{1-x_{2}} d x_{1}\left[V_{1}-T_{1}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
F_{12}= & \int_{0}^{1-x_{3}} d x_{1}\left[V_{1}-T_{1}+1 / 2 A_{1}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
F_{13}= & \int_{0}^{1-x_{2}} d x_{1}\left[V_{1}^{M}-T_{1}^{M}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
F_{14}= & \int_{0}^{1-x_{3}} d x_{1}\left[V_{1}^{M}-T_{1}^{M}+1 / 2 A_{1}^{M}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
F_{15}= & \int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{2} \int_{0}^{1-x_{2}} d x_{1}\left[T_{2}-T_{3}-T_{4}+T_{5}+T_{7}+T_{8}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
F_{16}= & \int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{3} \int_{0}^{1-x_{3}} d x_{1}\left[-T_{2}+T_{3}+T_{4}-T_{5}-T_{7}-T_{8}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
F_{17}= & \int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{2} \int_{0}^{1-x_{2}} d x_{1}\left[-T_{1}-T_{2}+2 T_{3}+2 T_{4}-T_{5}-T_{6}+2 A_{1}-2 A_{2}\right. \\
& \left.-2 A_{5}+2 A_{6}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right),
\end{aligned}
$$

$$
F_{18}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{3} \int_{0}^{1-x_{3}} d x_{1}\left[T_{1}-T_{2}-T_{5}+T_{6}-2 T_{7}-2 T_{8}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right)
$$

$$
F_{19}=\int_{0}^{1-x_{2}} d x_{1}\left[V_{1}^{M}-T_{1}^{M}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right)
$$

$$
F_{20}=\int_{0}^{1-x_{3}} d x_{1}\left[V_{1}^{M}\right]\left(x_{1}, 1-x_{1}-x_{2}, x_{3}\right)
$$

$$
F_{21}=\int_{0}^{1-x_{2}} d x_{1}\left[P_{1}-S_{1}+V_{1}+V_{2}-A_{1}+A_{2}-T_{3}-T_{7}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right)
$$

$$
\begin{aligned}
& F_{22}=\int_{0}^{1-x_{3}} d x_{1}\left[V_{3}-T_{1}-A_{3}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
& F_{23}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{2} \int_{0}^{1-x_{2}} d x_{1}\left[-T_{1}+T_{3}+T_{4}-T_{6}+T_{7}+T_{8}+A_{1}-A_{2}+A_{3}+A_{4}\right. \\
& \left.-A_{5}+A_{6}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{24}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{3} \int_{0}^{1-x_{3}} d x_{1}\left[T_{1}-T_{3}-T_{4}+T_{6}-T_{7}-T_{8}+A_{1}-A_{2}+A_{3}+A_{4}\right. \\
& \left.-A_{5}+A_{6}+V_{1}-V_{2}-V_{3}-V_{4}-V_{5}+V_{6}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right) \text {, } \\
& F_{25}=\int_{0}^{1-x_{2}} d x_{1}\left[S_{1}-P_{1}-V_{3}-2 V_{2}-A_{3}+T_{3}+T_{7}\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right),\right. \\
& F_{26}=\int_{0}^{1-x_{3}} d x_{1}\left[-V_{3}+T_{1}+T_{7}+A_{3}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
& F_{27}=\int_{0}^{1-x_{2}} d x_{1}\left[-2 V_{1}^{M}+T_{1}^{M}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{28}=\int_{0}^{1-x_{3}} d x_{1}\left[T_{1}^{M}\right]\left(x_{1}, 1-x_{1}-x_{2}, x_{3}\right), \\
& F_{29}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{2} \int_{0}^{1-x_{2}} d x_{1}\left[-T_{1}+T_{2}+T_{5}-T_{6}+2 T_{7}+2 T_{8}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{30}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{3} \int_{0}^{1-x_{3}} d x_{1}\left[T_{2}-T_{3}-T_{4}+T_{5}+T_{7}+T_{8}-V_{1}+V_{2}+V_{3}\right. \\
& \left.+V_{4}+V_{5}-V_{6}-A_{1}+A_{2}+A_{3}-A_{4}+A_{5}-A_{6}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
& F_{31}=\int_{\alpha}^{1} d x_{2} \int_{\alpha}^{1-x_{2}} d x_{1}\left[2 T_{1}-T_{4}-T_{7}-T_{8}+A_{3}-A_{4}-S_{1}+S_{2}\right. \\
& \left.+P_{1}-P_{2}-2 V_{1}+2 V_{2}+V_{3}+V_{4}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{32}=\int_{\alpha}^{1} d x_{3} \int_{\alpha}^{1-x_{3}} d x_{1}\left[T_{1}-T_{3}-T_{7}-V_{1}+V_{3}+V_{5}-A_{1}-A_{3}-A_{5}\right] \\
& \left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right),
\end{aligned}
$$

$$
\begin{aligned}
& F_{33}=\int_{0}^{1-x_{2}} d x_{1}\left[V_{1}-T_{1}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{34}=\int_{0}^{1-x_{3}} d x_{1}\left[V_{1}-T_{1}+A_{1}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
& F_{35}=\int_{0}^{1-x_{2}} d x_{1}\left[V_{1}^{M}-T_{1}^{M}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{36}=\int_{0}^{1-x_{3}} d x_{1}\left[V_{1}^{M}-T_{1}^{M}+A_{1}^{M}\right]\left(x_{1}, 1-x_{1}-x_{2}, x_{3}\right), \\
& F_{37}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{2} \int_{0}^{1-x_{2}} d x_{1}\left[T_{2}-T_{3}-T_{4}+T_{5}+T_{7}+T_{8}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{38}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{3} \int_{0}^{1-x_{3}} d x_{1}\left[T_{2}-T_{3}-T_{4}+T_{5}+T_{7}+T_{8}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
& F_{39}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{2} \int_{0}^{1-x_{2}} d x_{1}\left[T_{1}-T_{3}-T_{4}+T_{6}-T_{7}-T_{8}-A_{1}+A_{2}\right. \\
& \left.-A_{3}-A_{4}+A_{5}-A_{6}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right), \\
& F_{40}=\int_{0}^{\beta} d \alpha \int_{\alpha}^{1} d x_{3} \int_{0}^{1-x_{3}} d x_{1}\left[T_{1}-T_{3}-T_{4}+T_{6}-T_{7}-T_{8}+A_{1}-A_{2}+A_{3}\right. \\
& \left.+A_{4}-A_{5}+A_{6}-V_{1}+V_{2}+V_{3}+V_{4}+V_{5}-V_{6}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right), \\
& F_{41}=\int_{\alpha}^{1} d x_{2} \int_{\alpha}^{1-x_{2}} d x_{1}\left[A_{1}-A_{2}+A_{3}-V_{1}+V_{2}+V_{3}\right]\left(x_{1}, x_{2}, 1-x_{1}-x_{2}\right) \text {, } \\
& F_{42}=\int_{\alpha}^{1-x_{3}} d x_{1}\left[T_{1}-T_{3}+T_{7}\right]\left(x_{1}, 1-x_{1}-x_{3}, x_{3}\right) .
\end{aligned}
$$

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