

Search for jet extinction in the inclusive jet- p_T spectrum from proton-proton collisions at $\sqrt{s} = 8$ TeV

V. Khachatryan *et al.**

(CMS Collaboration)

(Received 29 May 2014; published 18 August 2014)

The first search at the LHC for the extinction of QCD jet production is presented, using data collected with the CMS detector corresponding to an integrated luminosity of 10.7 fb^{-1} of proton-proton collisions at a center-of-mass energy of 8 TeV. The extinction model studied in this analysis is motivated by the search for signatures of strong gravity at the TeV scale (terascale gravity) and assumes the existence of string couplings in the strong-coupling limit. In this limit, the string model predicts the suppression of all high-transverse-momentum standard model processes, including jet production, beyond a certain energy scale. To test this prediction, the measured transverse-momentum spectrum is compared to the theoretical prediction of the standard model. No significant deficit of events is found at high transverse momentum. A 95% confidence level lower limit of 3.3 TeV is set on the extinction mass scale.

DOI: [10.1103/PhysRevD.90.032005](https://doi.org/10.1103/PhysRevD.90.032005)

PACS numbers: 13.85.Rm, 13.87.-a, 14.70.Kv

I. INTRODUCTION

The scattering of high-energy particles in theories of quantum gravity is fundamentally different from that expected by the local quantum field theories of the standard model (SM) [1]. The Planck scale, the threshold at which quantum gravity becomes strong, is therefore a fundamental boundary beyond which some modification to the SM is required. The Planck scale differs from the electroweak scale by 16 orders of magnitude, creating what is commonly known as the hierarchy problem. There are many models that propose a mechanism by which these two scales are related to one another through the hypothesized existence of extra spatial dimensions. Propagation of gravitons through these extra dimensions could explain the relative weakness of gravity compared to the strong and electroweak interactions. Depending on the model, a variety of striking signatures of physics beyond the SM may be observed. As a result, models that predict terascale gravity have been the subject of numerous searches at the CERN LHC [2–11]. Some of these searches are designed to look for effects such as resonant production and decay of new states, e.g. Randall-Sundrum gravitons [12], as well as for continuum enhancements to SM processes from both virtual and direct graviton production [13]. Direct searches for production of microscopic black holes consider events with high transverse momentum (p_T) and multiple objects from the decay of possible high-entropy intermediate states [1,14,15].

As of yet, no signal indicative of terascale gravity has been found. Nevertheless, it has been suggested that evidence of terascale gravity could also be found through more subtle effects on the jet- p_T spectrum manifesting themselves as a deviation from the predictions of quantum chromodynamics (QCD) [1,14,16,17]. While the production of black holes or particles indicative of nonperturbative quantum gravity can have a rapidly increasing total cross section beyond some energy scale, their decay to isolated jets or other low-multiplicity final states could be suppressed, leading to a full suppression of high- p_T SM scattering processes (jet extinction). Because jet production is the leading SM process at high p_T , such effects would be initially noticeable as a jet extinction signature [17]. In this sense, the search for jet extinction is complementary to searches for black holes in high-multiplicity final states. These final states arise in the asymptotic limit, where black holes are expected to behave classically [15]. The extinction search explores an intermediate regime, where a high-multiplicity signature may not be readily observable.

There are several models that include extinction phenomena [16,17]. In this, the first search for extinction effects at the LHC, we consider a model with a large-width Veneziano form factor modification of QCD processes with an extinction mass scale M equivalent to the modified Planck scale [17]. This form factor is discussed in greater detail in Sec. 3. Beyond the scale M , the predominance of intermediate high-entropy string states will suppress high- p_T SM jet production. This search exploits techniques developed for the measurement of the differential jet production cross section as a function of p_T at the CMS [18] experiment to search for a modification of the jet- p_T spectrum consistent with extinction phenomena, in which there are fewer high- p_T jets than expected from the SM. This analysis is especially sensitive to the correlations of

* Full author list given at the end of the article.

Published by the American Physical Society under the terms of the Creative Commons Attribution 3.0 License. Further distribution of this work must maintain attribution to the author(s) and the published articles title, journal citation, and DOI.

the systematic uncertainties between bins in jet- p_T , so a detailed evaluation of the systematic uncertainties associated with the jet energy scale (JES) and the parton distribution functions (PDFs) is performed.

II. THE CMS DETECTOR

The central feature of the CMS detector [19] is a superconducting solenoid of 6 m internal diameter, providing a field of 3.8 T. Within the field volume are silicon pixel and strip trackers, a lead-tungstate crystal electromagnetic calorimeter (ECAL), and a brass and scintillator hadron calorimeter (HCAL). Muons are measured in gas-ionization detectors embedded in the steel flux-return yoke.

The CMS experiment uses a right-handed coordinate system, with the origin at the nominal interaction point, the x axis pointing to the center of the LHC ring, the y axis pointing up (perpendicular to the plane of the LHC ring), and the z axis along the counterclockwise-beam direction. The polar angle θ is measured from the positive z axis and the azimuthal angle ϕ is measured in the xy plane. The pseudorapidity is defined as $\eta = -\ln[\tan(\theta/2)]$.

The first level of the CMS trigger system is composed of customized hardware and uses information from the calorimeters and muon detectors to select events of interest within a 4 μ s interval following each beam crossing. The high-level trigger (HLT) [20] processor farm further decreases the event rate from about 100 kHz to about 400 Hz before the data are recorded for analysis.

III. MODELING OF THE SM AND EXTINCTION HYPOTHESES

The SM prediction for the jet- p_T spectrum is calculated at next-to-leading order (NLO) with the NLOJET++ program within the FASTNLO framework [21–23]. The CT10 PDF set [24] is used in this calculation. The renormalization and factorization scales, μ_R and μ_F , are set equal to the jet- p_T . The NLO jet spectra do not include nonperturbative (NP) effects or any modeling of the detector response. The NP effects, which account for hadronization and multiparton interactions, are incorporated as corrections determined from the PYTHIA 6.424 [25] Monte Carlo (MC) generator. The generator is used to simulate QCD events with and without NP effects. The corrections are derived from the ratio of the resulting p_T spectra. The NP correction decreases monotonically as a function of jet- p_T , from 1.03 at 592 GeV to 1.01 at 2500 GeV. This process is repeated using the HERWIG 2.4.2 [26] generator. The difference between the corrections derived from these generators is found to be negligible in the phase space of this analysis. The corrected NLO jet spectra are convolved with a function that models the jet energy resolution (JER) in the CMS detector [27]. These smeared spectra can be compared directly to the observed spectrum. The smeared NLO jet spectrum is referred to as $d\sigma^{\text{QCD}}/dp_{T,\text{NLO}}$.

This procedure is repeated to produce a smeared leading-order (LO) jet- p_T spectrum, labeled as $d\sigma^{\text{QCD}}/dp_{T,\text{LO}}$. The predicted spectrum does not include weak radiative corrections [28], but the impact of these corrections on our sensitivity to an extinction signature is evaluated during the limit-setting procedure.

The effects of extinction at LO are also modeled using the PYTHIA MC generator. The matrix elements of each color channel are modified by Veneziano-type form factors [17,29], which affect all $2 \rightarrow 2$ scattering amplitudes. The input parameters for these form factors are the extinction mass scale M and a dimensionless width parameter related to the strength of the string coupling. For small values of the width parameter, these form factors are similar to those that describe string resonances [29,30]. This is referred to as the weak-coupling limit. The regime where the width parameter is close to unity is known as the strong-coupling limit. In this limit, extinction physics rapidly overwhelms LO SM processes as well as any resonant string production. Beyond the scale M , scattering processes are dominated by a continuum of high-entropy intermediate states, which results in suppression of SM jet production [17]. This search assumes a width parameter of one, the absolute strong-coupling limit of the string model. Values of the width above one represent a very different phenomenology where the form factors no longer monotonically decrease as a function of jet momentum. This range of the width parameter has not been studied in this analysis.

The effects of extinction are predominantly found in $2 \rightarrow 2$ scattering processes. Such processes are dominated by the LO calculation at a given p_T scale. The signal is approximated with a LO generator. The extinction process is assumed to have a very weak effect on higher-order interactions. A sigmoid function provides a good functional fit of the effect of the Veneziano form factors on the LO jet- p_T spectrum [17]:

$$F(p_T, M) = \frac{1}{1 + \exp\left(\frac{p_T - p_{T,1/2}(M)}{p_{T,0}(M)}\right)}. \quad (1)$$

Here, $p_{T,1/2}$ describes the p_T threshold at which LO jet production is reduced to half the SM expectation, while $p_{T,0}$ indicates how quickly the LO cross section exponentially falls relative to the SM prediction. This relation yields the following equation for the jet- p_T spectrum assuming extinction at LO, where σ^{Ext} is the jet production cross section assuming extinction:

$$\frac{d\sigma^{\text{Ext}}}{dp_{T,\text{LO}}} = \frac{d\sigma^{\text{QCD}}}{dp_{T,\text{LO}}} F(p_T, M) \quad (2)$$

and at NLO:

$$\frac{d\sigma^{\text{Ext}}}{dp_{T,\text{NLO}}} = \frac{d\sigma^{\text{QCD}}}{dp_{T,\text{NLO}}} - \frac{d\sigma^{\text{QCD}}}{dp_{T,\text{LO}}} + \frac{d\sigma^{\text{Ext}}}{dp_{T,\text{LO}}}. \quad (3)$$

Several simulations of LO jet production are performed, assuming values of M between 2 and 5 TeV in increments of 500 GeV. The jet- p_T spectrum is produced at NLO for each sample using NP corrections and resolution smearing as described above. The values of $p_{T,1/2}(M)$ and $p_{T,0}(M)$ are extracted from a fit of $F(p_T, M)$ to the expected p_T distribution for each value of M . The intermediate values of $p_{T,1/2}(M)$ and $p_{T,0}(M)$ are interpolated between these fitted points. The fitted value of $p_{T,0}(M)$ is nearly independent of M and ranges between 260 and 330 GeV, while $p_{T,1/2}(M)$ is about half of M . The systematic uncertainty associated with the choice of fit is negligible.

For finite values of M , the predicted jet- p_T spectrum is suppressed in systems with an invariant mass above M . At very large values of M , the SM and extinction spectra become identical.

IV. EVENT RECONSTRUCTION AND SELECTION

A particle-flow algorithm [31,32] is used to reconstruct the events. Jets are formed by clustering the reconstructed particle-flow objects using the anti- k_T algorithm [33] with a distance parameter R of 0.7. This value is larger than the usual distance parameter of 0.5 used in most CMS analyses. The larger cluster size reduces the likelihood that jets will be lost because of detector effects. The jet transverse momentum resolution is typically 15% at $p_T = 10$ GeV, 8% at 100 GeV, and 4% at 1 TeV. Jet energy corrections are derived from simulation and are confirmed with measurements of energy balance in recorded dijet and photon + jet events. The combined corrections are approximately 5%–10%, depending on the pseudorapidity and p_T of the jet. To suppress spurious signals from detector noise [34], jets are required to satisfy stringent selection criteria [35]. Specifically, each jet must contain at least two particles, one of which is a charged hadron. Additionally, each of the jet energy fractions carried by neutral hadrons, photons, electrons, and muons must be less than 90%. This analysis is conducted in a regime where the purity and acceptance of the jets in data are both close to unity, and therefore no systematic uncertainty is attributed to the selection criteria.

The data used in this analysis were collected from an HLT trigger that accepted events containing at least one jet with $p_T > 320$ GeV. An offset is applied to trigger-selected jets to subtract the energy deposited as a result of additional interactions per beam crossing (pileup); this offset does not affect the trigger efficiency. Events with objects originating from an interaction within an LHC beam crossing are selected by requiring the presence of at least one primary vertex within 24 cm of the detector center along the z axis. The primary event vertex is chosen from all reconstructed vertices by selecting the one with the largest sum of the p_T^2 of all associated tracks. For the purpose of additional noise suppression, the missing transverse energy, defined as the magnitude of the vector sum p_T of all reconstructed particle-flow objects, must be less than

30% of the total transverse energy deposited in the detector. All jets in each event that pass the selection criteria are binned as a function of jet- p_T , following a convention adopted by other inclusive-jet analyses in CMS. The bin widths are variable, increasing with jet- p_T and corresponding approximately to the jet- p_T resolution [18]. Jets are required to have $p_T > 592$ GeV and pseudorapidity $|\eta| < 1.5$ to ensure that the trigger is at least 99% efficient in all p_T bins used. This search is performed in 18 p_T bins between 592 and 2500 GeV.

A comparison between the observed inclusive jet- p_T spectrum and the spectrum predicted at NLO with the CT10 PDF set is shown in Figs. 1 and 2. The predicted spectrum includes nonperturbative corrections and smearing by the detector response and is normalized to the total number of jets in data that pass all selection criteria. However, in the comparison of the model to the data as described in Sec. 5, the SM distribution is instead normalized to the number of jets expected given an integrated luminosity of 10.7 fb^{-1} . The number of jets observed in data is 3% lower than the number expected assuming the CT10 PDF set at NLO. This discrepancy is attributed to uncertainty in the PDF parameters, scale variations in the cross section calculation, or

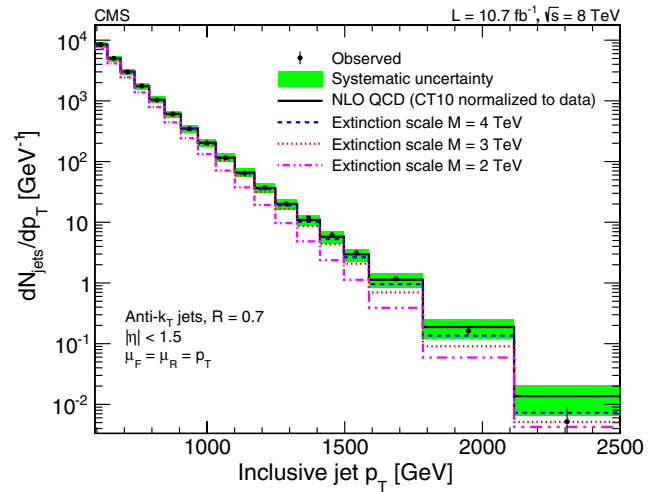


FIG. 1 (color online). Inclusive jet- p_T spectrum (points) for $|\eta| < 1.5$, as observed in data. The SM NLO simulation with nonperturbative corrections, convolved with the detector response and normalized to the total number of jets observed in data, is shown by the solid line. The spectra predicted by the extinction model are defined relative to the SM prediction as described by Eq. (3) for the values of $M = 2, 3$, and 4 TeV and shown by the dashed lines. The colored band shows the magnitude of the sources of systematic uncertainty added in quadrature. These sources include the JES, JER, PDFs, and scale variations. An additional source of systematic uncertainty is attributed to the integrated luminosity during all formal comparisons between the data and models but has little impact on the sensitivity to an extinction signature. The renormalization scale (μ_R) and factorization scale (μ_F) are set to the p_T of the hard-scattered parton.

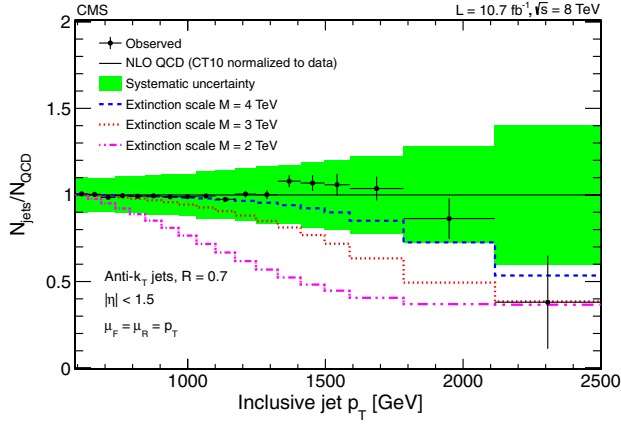


FIG. 2 (color online). The ratio of the inclusive jet- p_T spectrum to the NLO QCD prediction with nonperturbative corrections and convolved with the detector resolution. The horizontal bars on the data indicate the width of each bin in p_T . The colored band shows the quadratic sum of the sources of systematic uncertainty, including JES, JER, PDFs, and scale variations. The uncertainty in the integrated luminosity is excluded, as the model predictions have been normalized to the number of jets observed in data. The dashed lines indicate the effects of extinction at three different values of the extinction mass scale, $M = 2, 3$, and 4 TeV.

uncertainty in the total integrated luminosity. As the search for an extinction signature is only concerned with the shape of the jet- p_T spectrum, a small shift in the absolute normalization has little impact on the sensitivity. In Figs. 1 and 2 the data and the extinction model are compared after any differences in the normalization have been resolved. In these figures, the quadratic sum of all sources of systematic uncertainty is shown. The total systematic uncertainty includes contributions from both theoretical and experimental sources. The theoretical uncertainty is composed of the uncertainty from the PDFs as well as the uncertainty obtained by varying the renormalization and factorization scales. The experimental uncertainty is derived from the uncertainties in the JES and JER. During the formal comparison of the model to data where the predicted spectrum is not normalized to the number of jets observed, an additional source of uncertainty is attributed to the integrated luminosity. Figure 2 shows the ratio of the inclusive spectrum to the SM NLO expectation and includes the predicted spectra from the extinction model for three different values of the extinction mass scale M .

V. STATISTICAL METHOD AND SYSTEMATIC UNCERTAINTIES

To distinguish between SM NLO jet production and the alternative hypothesis (jet extinction), a profile-likelihood ratio test statistic [36] is constructed as a function of a signal strength parameter, $\beta \equiv M^{-2}$. The variable β is chosen so that as $\beta \rightarrow 0$ the extinction model approaches the SM prediction.

We set limits using the modified-frequentist criterion CL_s [37,38]. All sources of systematic uncertainty are treated as nuisance parameters with log-normal prior constraints and are constructed in the likelihood to have the same value across all jet- p_T bins. This construction implicitly assumes that the systematic uncertainties are completely correlated in jet- p_T .

To account for correlations in the JES and PDF uncertainties between p_T bins, the uncertainties are subdivided into their underlying components. These individual components are strongly correlated across all p_T bins and tend to be dominant at different values of jet- p_T . As an example, uncertainties in the gluon PDF will be dominant at low p_T compared to uncertainties in the quark PDFs. The JES uncertainty is decomposed into each of its orthogonal sources. For the PDF uncertainty, the contributions from each of the eigenvectors in the CT10 [24] PDF set are evaluated separately. As a cross-check, the search is repeated with respect to the MSTW2008 [39] PDF set. Among the PDF sets in common use, the CT10 set predicts the highest inclusive jet cross section at high p_T , while the MSTW2008 set gives one of the lowest. The results derived with respect to these two PDF sets serve as bounds on the result expected when using other sets, including those which are used in comparison to dedicated measurements of the inclusive jet production cross section [18], such as NNPDF [40], HERA [41], or ABKM [42].

The CT10 PDF set comprises a central prediction and 26 eigenvectors. The central prediction assumes all PDF input parameters are set to their central values. Each eigenvector pair corresponds to the upward and downward uncertainty in one of those input parameters. The difference between the predictions of each eigenvector pair and the central prediction is taken as a source of systematic uncertainty at $\pm 1\sigma$. A source of systematic uncertainty is defined as nontrivial if, at one standard deviation in either direction, it produces a shift in any p_T bin greater than 1% of the occupancy given by the central prediction. Under this definition, 15 of the 26 CT10 eigenvectors are found to be nontrivial.

The relative uncertainty described by the combined variation of these eigenvector sets in quadrature and the scale variations are shown in Fig. 3 as a function of jet- p_T . The uncertainties associated with the renormalization and factorization scales are computed by varying the scales coherently up and down by a factor of 2. As the effect of extinction on the jet- p_T spectrum is expressed relative to the SM prediction, by construction the PDF variations do not affect any of the extinction parameters.

Given the exponentially falling nature of the inclusive jet- p_T spectrum, the JES is one of the dominant sources of systematic uncertainty. The JES uncertainty is composed of 19 orthogonal sources. Of these, seven are found to be nontrivial according to the criterion defined above: the absolute p_T scale; the single pion response in the ECAL;

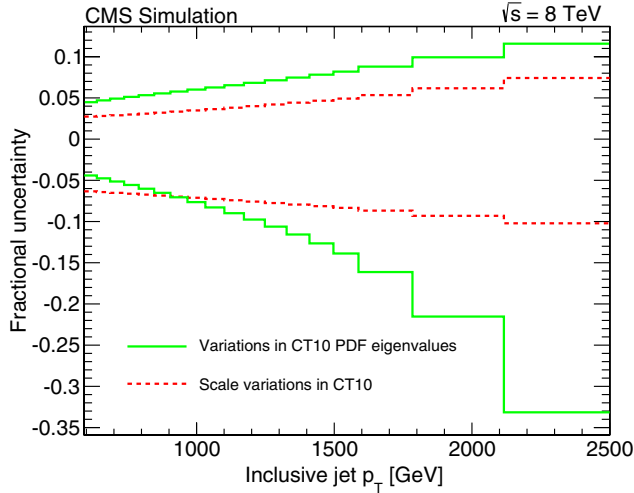


FIG. 3 (color online). Uncertainty at ± 1 standard deviation described by the combined variations of all CT10 PDF eigenvectors added in quadrature (solid lines), as well as the scale variations (dotted lines). The uncertainty is expressed as a fraction of the central occupancy of each p_T bin. For the fit of the model to data and the limit-setting procedure, the PDF uncertainty is subdivided into individual sources for each eigenvector pair.

the single pion response in the HCAL; the flavor composition correction; the time dependence; the pileup p_T scale; and the extrapolation of the absolute scale into the high- p_T regime [27]. The effects of JER are also included as nuisance parameters. The uncertainty in luminosity is taken as a constant scale factor with a 2.6% relative uncertainty

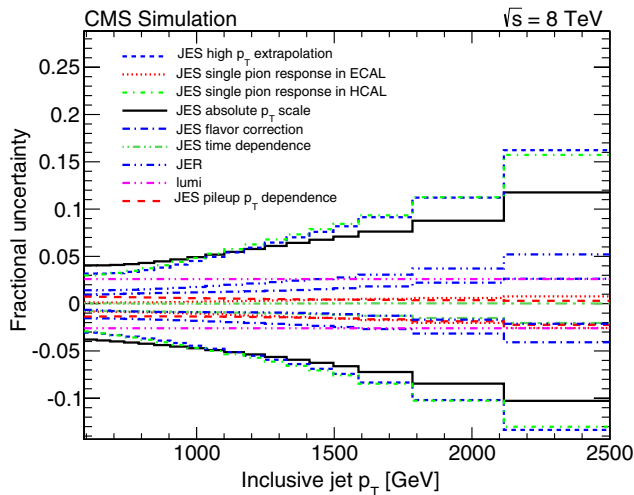


FIG. 4 (color online). Systematic uncertainty from all experimental sources at ± 1 standard deviation, expressed as a fraction of the central occupancy of each p_T bin. The luminosity uncertainty is constant in jet- p_T , while the JES and JER uncertainties are modeled as transfer matrices between all p_T bins. The seven nontrivial sources of JES uncertainty are shown (out of 19 total).

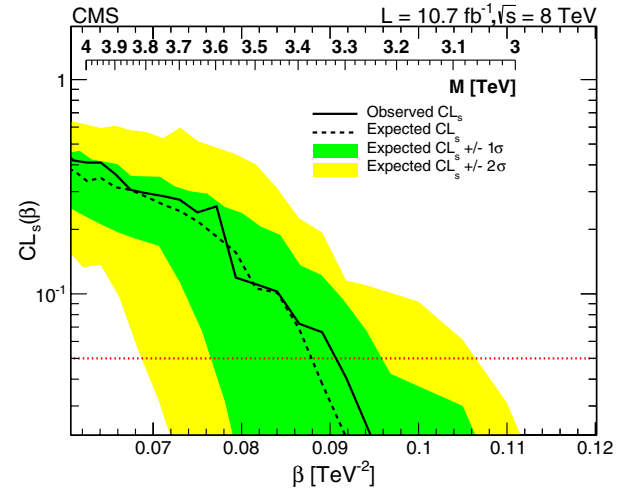


FIG. 5 (color online). The results of a CL_s scan in the extinction mass scale, $\beta = M^{-2}$. The observed dependence of CL_s on β is shown by the solid line. The observed upper limit on β is 0.090 TeV^{-2} at 95% C.L. (indicated by the horizontal dotted line), corresponding to a lower limit of 3.3 TeV on the extinction mass scale M . The dashed line indicates the expected median of results for the SM hypothesis, while the green (dark) and yellow (light) bands indicate the quantiles, which contain 68% and 95% of the expected results, respectively.

[43]. The relative uncertainty of all nontrivial detector-related sources of systematic uncertainty (JES, JER, and integrated luminosity) is shown in Fig. 4 as a function of jet- p_T .

Including systematic uncertainties, the best-fit value of β is $(0.008 \pm 0.033) \text{ TeV}^{-2}$, which is consistent with the SM expectation.

The dependence of CL_s on the parameter β is shown in Fig. 5. The observed upper limit on β is 0.090 TeV^{-2} at 95% confidence level (C.L.), translating to an observed lower limit on M of 3.3 TeV. The expected upper limit on β is 0.088 TeV^{-2} at 95% C.L., corresponding to an expected lower limit on M of 3.4 TeV. These relatively close expected and observed values reflect good agreement between the observed data and the null hypothesis.

As an additional check, the limit-setting procedure is repeated using the MSTW2008 PDF set [39] to derive the SM hypothesis. The limits obtained using the CT10 and MSTW2008 PDFs agree to within 10%. As the MSTW2008 PDFs predict a lower cross section at very high jet- p_T compared to CT10, the limit produced in this check is less conservative.

Finally, the limits have been calculated including weak radiative corrections to the SM prediction, with a decrease of less than 100 GeV to the exclusion region.

VI. SUMMARY

The first search for the extinction of jet production has been performed at the LHC using proton-proton collision

data at $\sqrt{s} = 8$ TeV collected by the CMS detector and corresponding to an integrated luminosity of 10.7 fb^{-1} . The extinction model studied in this analysis is motivated by the search for signatures of terascale gravity at the LHC and assumes the existence of string couplings in the strong-coupling limit. In this limit, the string model predicts suppression of high- p_T jet production beyond an extinction mass scale M . A detailed comparison between the measured p_T spectrum and the theoretical prediction is conducted. No significant deficit of events is found at high transverse momentum. A 95% confidence level lower limit of 3.3 TeV is set on the extinction mass scale M .

ACKNOWLEDGMENTS

We congratulate our colleagues in the CERN accelerator departments for the excellent performance of the LHC and thank the technical and administrative staffs at CERN and at other CMS institutes for their contributions to the success of the CMS effort. In addition, we gratefully acknowledge the computing centers and personnel of the Worldwide LHC Computing Grid for delivering so effectively the computing infrastructure essential to our analyses. Finally, we acknowledge the enduring support for the construction and operation of the LHC and the CMS detector provided by the following funding agencies: BMWF and FWF (Austria); FNRS and FWO (Belgium); CNPq, CAPES, FAPERJ, and FAPESP (Brazil); MES (Bulgaria); CERN; CAS, MoST, and NSFC (China); COLCIENCIAS (Colombia); MSES and CSF (Croatia); RPF (Cyprus); MoER, ERC IUT and ERDF (Estonia); Academy of

Finland, MEC, and HIP (Finland); CEA and CNRS/IN2P3 (France); BMBF, DFG, and HGF (Germany); GSRT (Greece); OTKA and NIH (Hungary); DAE and DST (India); IPM (Iran); SFI (Ireland); INFN (Italy); NRF and WCU (Republic of Korea); LAS (Lithuania); MOE and UM (Malaysia); CINVESTAV, CONACYT, SEP, and UASLP-FAI (Mexico); MBIE (New Zealand); PAEC (Pakistan); MSHE and NSC (Poland); FCT (Portugal); JINR (Dubna); MON, RosAtom, RAS and RFBR (Russia); MESTD (Serbia); SEIDI and CPAN (Spain); Swiss Funding Agencies (Switzerland); MST (Taipei); ThEPCenter, IPST, STAR and NSTDA (Thailand); TUBITAK and TAEK (Turkey); NASU and SFFR (Ukraine); STFC (United Kingdom); DOE and NSF (USA). Individuals have received support from the Marie-Curie program and the European Research Council and EPLANET (European Union); the Leventis Foundation; the A. P. Sloan Foundation; the Alexander von Humboldt Foundation; the Belgian Federal Science Policy Office; the Fonds pour la Formation à la Recherche dans l'Industrie et dans l'Agriculture (FRIA-Belgium); the Agentschap voor Innovatie door Wetenschap en Technologie (IWT-Belgium); the Ministry of Education, Youth and Sports (MEYS) of the Czech Republic; the Council of Science and Industrial Research, India; the HOMING PLUS program of Foundation for Polish Science, cofinanced from European Union, Regional Development Fund; the Compagnia di San Paolo (Torino); and the Thalys and Aristeia programs cofinanced by EU-ESF and the Greek NSRF.

-
- [1] T. Banks and W. Fischler, [arXiv:hep-th/9906038](#).
 - [2] CMS Collaboration, *J. High Energy Phys.* **04** (2012) 061.
 - [3] CMS Collaboration, *J. High Energy Phys.* **07** (2013) 178.
 - [4] CMS Collaboration, *Phys. Rev. Lett.* **108**, 111801 (2012).
 - [5] CMS Collaboration, *Phys. Lett. B* **704**, 123 (2011).
 - [6] CMS Collaboration, *J. High Energy Phys.* **05** (2011) 093.
 - [7] ATLAS Collaboration, *Phys. Lett. B* **709**, 322 (2012).
 - [8] ATLAS Collaboration, *Phys. Lett. B* **710**, 538 (2012).
 - [9] ATLAS Collaboration, *Phys. Lett. B* **712**, 331 (2012).
 - [10] CMS Collaboration, *J. High Energy Phys.* **09** (2012) 094.
 - [11] CMS Collaboration, *Phys. Rev. Lett.* **108**, 261803 (2012).
 - [12] L. Randall and R. Sundrum, *Phys. Rev. Lett.* **83**, 3370 (1999).
 - [13] N. Arkani-Hamed, S. Dimopoulos, and G. Dvali, *Phys. Rev. D* **59**, 086004 (1999).
 - [14] S. B. Giddings and S. D. Thomas, *Phys. Rev. D* **65**, 056010 (2002).
 - [15] S. Dimopoulos and G. Landsberg, *Phys. Rev. Lett.* **87**, 161602 (2001).
 - [16] P. Meade and L. Randall, *J. High Energy Phys.* **05** (2008) 003.
 - [17] C. Kilic, A. Lath, K. Rose, and S. Thomas, *Phys. Rev. D* **89**, 016003 (2014).
 - [18] CMS Collaboration, *Phys. Rev. Lett.* **107**, 132001 (2011).
 - [19] CMS Collaboration, *J. Instrum.* **3**, S08004 (2008).
 - [20] CMS Collaboration, *Eur. Phys. J. C* **46**, 605 (2006).
 - [21] Z. Nagy, *Phys. Rev. Lett.* **88**, 122003 (2002).
 - [22] Z. Nagy, *Phys. Rev. D* **68**, 094002 (2003).
 - [23] D. Britzger, K. Rabbertz, F. Stober, and M. Wobisch (unpublished).
 - [24] P. M. Nadolsky, H.-L. Lai, Q.-H. Cao, J. Huston, J. Pumplin, D. Stump, W.-K. Tung, and C.-P. Yuan, *Phys. Rev. D* **78**, 013004 (2008).
 - [25] T. Sjöstrand, S. Mrenna, and P. Z. Skands, *J. High Energy Phys.* **05** (2006) 026.
 - [26] M. Bähr, S. Gieseke, M. A. Gigg, D. Grellscheid, K. Hamilton, O. Latunde-Dada, S. Plätzer, P. Richardson, M. H. Seymour, A. Sherstnev, and B. R. Webber, *Eur. Phys. J. C* **58**, 639 (2008).
 - [27] CMS Collaboration, *J. Instrum.* **6**, P11002 (2011).
 - [28] S. Dittmaier, A. Huss, and C. Speckner, *J. High Energy Phys.* **11** (2012) 095.
 - [29] D. Lust, S. Stieberger, and T. R. Taylor, *Nucl. Phys.* **B808**, 1 (2009).

- [30] E. A. Mirabelli, M. Perelstein, and M. E. Peskin, *Phys. Rev. Lett.* **82**, 2236 (1999).
- [31] CMS Collaboration, CMS Physics Analysis Summary No. CMS-PAS-PFT-09-001, 2009, <http://cdsweb.cern.ch/record/1194487>.
- [32] CMS Collaboration, CMS Physics Analysis Summary No. CMS-PAS-PFT-10-001, 2010, <http://cdsweb.cern.ch/record/1247373>.
- [33] M. Cacciari, G. P. Salam, and G. Soyez, *J. High Energy Phys.* **04** (2008) 063.
- [34] CMS Collaboration, *J. Instrum.* **5**, T03014 (2010).
- [35] CMS Collaboration, CMS Physics Analysis Summary No. CMS-PAS-JME-09-008, 2010, <http://cdsweb.cern.ch/record/1259924>.
- [36] G. Cowan, K. Cranmer, E. Gross, and O. Vitells, *Eur. Phys. J. C* **71**, 1554 (2011).
- [37] T. Junk, *Nucl. Instrum. Methods Phys. Res., Sect. A* **434**, 435 (1999).
- [38] A. L. Read, *J. Phys. G* **28**, 2693 (2002).
- [39] A. D. Martin, W. J. Stirling, R. S. Thorne, and G. Watt, *Eur. Phys. J. C* **63**, 189 (2009).
- [40] R. D. Ball, L. Del Debbio, S. Forte, A. Guffanti, J. I. Latorre, J. Rojo, and M. Ubiali (NNPDF Collaboration), *Nucl. Phys.* **B838**, 136 (2010).
- [41] F. D. Aaron *et al.* (H1 and ZEUS Collaboration), *J. High Energy Phys.* **01** (2010) 109.
- [42] S. Alekhin, J. Blümlein, S. Klein, and S. Moch, *Phys. Rev. D* **81**, 014032 (2010).
- [43] CMS Collaboration, CMS Physics Analysis Summary No. CMS-PAS-LUM-13-001, 2012, <http://cdsweb.cern.ch/record/1598864>.

V. Khachatryan,¹ A. M. Sirunyan,¹ A. Tumasyan,¹ W. Adam,² T. Bergauer,² M. Dragicevic,² J. Erö,² C. Fabjan,^{2,b} M. Friedl,² R. Frühwirth,^{2,b} V. M. Ghete,² C. Hartl,² N. Hörmann,² J. Hrubec,² M. Jeitler,^{2,b} W. Kiesenhofer,² V. Knünz,² M. Krammer,^{2,b} I. Krätschmer,² D. Liko,² I. Mikulec,² D. Rabadý,^{2,c} B. Rahbaran,² H. Rohringer,² R. Schöfbeck,² J. Strauss,² A. Taurok,² W. Treberer-Treberspurg,² W. Waltenberger,² C.-E. Wulz,^{2,b} V. Mossolov,³ N. Shumeiko,³ J. Suarez Gonzalez,³ S. Alderweireldt,⁴ M. Bansal,⁴ S. Bansal,⁴ T. Cornelis,⁴ E. A. De Wolf,⁴ X. Janssen,⁴ A. Knutsson,⁴ S. Luyckx,⁴ S. Ochesanu,⁴ B. Roland,⁴ R. Rougny,⁴ M. Van De Klundert,⁴ H. Van Haevermaet,⁴ P. Van Mechelen,⁴ N. Van Remortel,⁴ A. Van Spilbeeck,⁴ F. Blekman,⁵ S. Blyweert,⁵ J. D'Hondt,⁵ N. Daci,⁵ N. Heracleous,⁵ A. Kalogeropoulos,⁵ J. Keaveney,⁵ T. J. Kim,⁵ S. Lowette,⁵ M. Maes,⁵ A. Olbrechts,⁵ Q. Python,⁵ D. Strom,⁵ S. Tavernier,⁵ W. Van Doninck,⁵ P. Van Mulders,⁵ G. P. Van Onsem,⁵ I. Vilella,⁵ C. Caillol,⁶ B. Clerbaux,⁶ G. De Lentdecker,⁶ D. Dobur,⁶ L. Favart,⁶ A. P. R. Gay,⁶ A. Grebenyuk,⁶ A. Léonard,⁶ A. Mohammadi,⁶ L. Perniè,^{6,c} T. Reis,⁶ T. Seva,⁶ L. Thomas,⁶ C. Vander Velde,⁶ P. Vanlaer,⁶ J. Wang,⁶ V. Adler,⁷ K. Beernaert,⁷ L. Benucci,⁷ A. Cimmino,⁷ S. Costantini,⁷ S. Crucy,⁷ S. Dildick,⁷ A. Fagot,⁷ G. Garcia,⁷ B. Klein,⁷ J. Mccartin,⁷ A. A. Ocampo Rios,⁷ D. Ryckbosch,⁷ S. Salva Diblen,⁷ M. Sigamani,⁷ N. Strobbe,⁷ F. Thyssen,⁷ M. Tytgat,⁷ E. Yazgan,⁷ N. Zaganidis,⁷ S. Basegmez,⁸ C. Beluffi,^{8,d} G. Bruno,⁸ R. Castello,⁸ A. Caudron,⁸ L. Ceard,⁸ G. G. Da Silva,⁸ C. Delaere,⁸ T. du Pree,⁸ D. Favart,⁸ L. Forthomme,⁸ A. Giammanco,^{8,e} J. Hollar,⁸ P. Jez,⁸ M. Komm,⁸ V. Lemaître,⁸ J. Liao,⁸ C. Nuttens,⁸ D. Pagano,⁸ A. Pin,⁸ K. Piotrkowski,⁸ A. Popov,^{8,f} L. Quertenmont,⁸ M. Selvaggi,⁸ M. Vidal Marono,⁸ J. M. Vizan Garcia,⁸ N. Belyi,⁹ T. Caeberts,⁹ E. Daubie,⁹ G. H. Hammad,⁹ G. A. Alves,¹⁰ M. Correa Martins Junior,¹⁰ T. Dos Reis Martins,¹⁰ M. E. Pol,¹⁰ W. L. Aldá Júnior,¹¹ W. Carvalho,¹¹ J. Chinellato,^{11,g} A. Custódio,¹¹ E. M. Da Costa,¹¹ D. De Jesus Damiao,¹¹ C. De Oliveira Martins,¹¹ S. Fonseca De Souza,¹¹ H. Malbouisson,¹¹ M. Malek,¹¹ D. Matos Figueiredo,¹¹ L. Mundim,¹¹ H. Nogima,¹¹ W. L. Prado Da Silva,¹¹ J. Santaolalla,¹¹ A. Santoro,¹¹ A. Sznajder,¹¹ E. J. Tonelli Manganote,^{11,g} A. Vilela Pereira,¹¹ C. A. Bernardes,^{12b} F. A. Dias,^{12a,h} T. R. Fernandez Perez Tomei,^{12a} E. M. Gregores,^{12b} P. G. Mercadante,^{12b} S. F. Novaes,^{12a} S. S. Padula,^{12a} A. Aleksandrov,¹³ V. Genchev,^{13,c} P. Iaydjiev,¹³ A. Marinov,¹³ S. Piperov,¹³ M. Rodozov,¹³ G. Sultanov,¹³ M. Vutova,¹³ A. Dimitrov,¹⁴ I. Glushkov,¹⁴ R. Hadjiiska,¹⁴ V. Kozhuharov,¹⁴ L. Litov,¹⁴ B. Pavlov,¹⁴ P. Petkov,¹⁴ J. G. Bian,¹⁵ G. M. Chen,¹⁵ H. S. Chen,¹⁵ M. Chen,¹⁵ R. Du,¹⁵ C. H. Jiang,¹⁵ D. Liang,¹⁵ S. Liang,¹⁵ R. Plestina,^{15,i} J. Tao,¹⁵ X. Wang,¹⁵ Z. Wang,¹⁵ C. Asawatangtrakuldee,¹⁶ Y. Ban,¹⁶ Y. Guo,¹⁶ Q. Li,¹⁶ W. Li,¹⁶ S. Liu,¹⁶ Y. Mao,¹⁶ S. J. Qian,¹⁶ D. Wang,¹⁶ L. Zhang,¹⁶ W. Zou,¹⁶ C. Avila,¹⁷ L. F. Chaparro Sierra,¹⁷ C. Florez,¹⁷ J. P. Gomez,¹⁷ B. Gomez Moreno,¹⁷ J. C. Sanabria,¹⁷ N. Godinovic,¹⁸ D. Lelas,¹⁸ D. Polic,¹⁸ I. Puljak,¹⁸ Z. Antunovic,¹⁹ M. Kovac,¹⁹ V. Brigljevic,²⁰ K. Kadija,²⁰ J. Luetic,²⁰ D. Mekterovic,²⁰ L. Sudic,²⁰ A. Attikis,²¹ G. Mavromanolakis,²¹ J. Mousa,²¹ C. Nicolaou,²¹ F. Ptochos,²¹ P. A. Razis,²¹ M. Bodlak,²² M. Finger,²² M. Finger Jr.,²² Y. Assran,^{23,j} S. Elgammal,^{23,k} M. A. Mahmoud,^{23,l} A. Radi,^{23,k,m} M. Kadastik,²⁴ M. Murumaa,²⁴ M. Raidal,²⁴ A. Tiko,²⁴ P. Eerola,²⁵ G. Fedi,²⁵ M. Voutilainen,²⁵ J. Härkönen,²⁶ V. Karimäki,²⁶ R. Kinnunen,²⁶ M. J. Kortelainen,²⁶ T. Lampén,²⁶ K. Lassila-Perini,²⁶ S. Lehti,²⁶ T. Lindén,²⁶ P. Luukka,²⁶ T. Mäenpää,²⁶ T. Peltola,²⁶ E. Tuominen,²⁶ J. Tuominiemi,²⁶ E. Tuovinen,²⁶ L. Wendland,²⁶ T. Tuuva,²⁷ M. Besancon,²⁸ F. Couderc,²⁸ M. Dejardin,²⁸ D. Denegri,²⁸ B. Fabbro,²⁸ J. L. Faure,²⁸ C. Favaro,²⁸ F. Ferri,²⁸ S. Ganjour,²⁸ A. Givernaud,²⁸

P. Gras,²⁸ G. Hamel de Monchenault,²⁸ P. Jarry,²⁸ E. Locci,²⁸ J. Malcles,²⁸ A. Nayak,²⁸ J. Rander,²⁸ A. Rosowsky,²⁸ M. Titov,²⁸ S. Baffioni,²⁹ F. Beaudette,²⁹ P. Busson,²⁹ C. Charlot,²⁹ T. Dahms,²⁹ M. Dalchenko,²⁹ L. Dobrzynski,²⁹ N. Filipovic,²⁹ A. Florent,²⁹ R. Granier de Cassagnac,²⁹ L. Mastrolorenzo,²⁹ P. Miné,²⁹ C. Mironov,²⁹ I. N. Naranjo,²⁹ M. Nguyen,²⁹ C. Ochando,²⁹ P. Paganini,²⁹ R. Salerno,²⁹ J. B. Sauvan,²⁹ Y. Sirois,²⁹ C. Veelken,²⁹ Y. Yilmaz,²⁹ A. Zabi,²⁹ J.-L. Agram,^{30,n} J. Andrea,³⁰ A. Aubin,³⁰ D. Bloch,³⁰ J.-M. Brom,³⁰ E. C. Chabert,³⁰ C. Collard,³⁰ E. Conte,^{30,n} J.-C. Fontaine,^{30,n} D. Gelé,³⁰ U. Goerlach,³⁰ C. Goetzmann,³⁰ A.-C. Le Bihan,³⁰ P. Van Hove,³⁰ S. Gadrat,³¹ S. Beauceron,³² N. Beaupere,³² G. Boudoul,^{32,c} S. Brochet,³² C. A. Carrillo Montoya,³² J. Chasserat,³² R. Chierici,³² D. Contardo,^{32,c} P. Depasse,³² H. El Mamouni,³² J. Fan,³² J. Fay,³² S. Gascon,³² M. Gouzevitch,³² B. Ille,³² T. Kurca,³² M. Lethuillier,³² L. Mirabito,³² S. Perries,³² J. D. Ruiz Alvarez,³² D. Sabes,³² L. Sgandurra,³² V. Sordini,³² M. Vander Donckt,³² P. Verdier,³² S. Viret,³² H. Xiao,³² Z. Tsamalaidze,^{33,o} C. Autermann,³⁴ S. Beranek,³⁴ M. Bontenackels,³⁴ B. Calpas,³⁴ M. Edelhoff,³⁴ L. Feld,³⁴ O. Hindrichs,³⁴ K. Klein,³⁴ A. Ostapchuk,³⁴ A. Perieanu,³⁴ F. Raupach,³⁴ J. Sammet,³⁴ S. Schael,³⁴ D. Sprenger,³⁴ H. Weber,³⁴ B. Wittmer,³⁴ V. Zhukov,^{34,f} M. Ata,³⁵ J. Caudron,³⁵ E. Dietz-Laursonn,³⁵ D. Duchardt,³⁵ M. Erdmann,³⁵ R. Fischer,³⁵ A. Güth,³⁵ T. Hebbeker,³⁵ C. Heidemann,³⁵ K. Hoepfner,³⁵ D. Klingebiel,³⁵ S. Knutzen,³⁵ P. Kreuzer,³⁵ M. Merschmeyer,³⁵ A. Meyer,³⁵ M. Olschewski,³⁵ K. Padeken,³⁵ P. Papacz,³⁵ H. Reithler,³⁵ S. A. Schmitz,³⁵ L. Sonnenschein,³⁵ D. Teyssier,³⁵ S. Thüer,³⁵ M. Weber,³⁵ V. Cherepanov,³⁶ Y. Erdogan,³⁶ G. Flügge,³⁶ H. Geenen,³⁶ M. Geisler,³⁶ W. Haj Ahmad,³⁶ F. Hoehle,³⁶ B. Kargoll,³⁶ T. Kress,³⁶ Y. Kuessel,³⁶ J. Lingemann,^{36,c} A. Nowack,³⁶ I. M. Nugent,³⁶ L. Perchalla,³⁶ O. Pooth,³⁶ A. Stahl,³⁶ I. Asin,³⁷ N. Bartosik,³⁷ J. Behr,³⁷ W. Behrenhoff,³⁷ U. Behrens,³⁷ A. J. Bell,³⁷ M. Bergholz,^{37,p} A. Bethani,³⁷ K. Borras,³⁷ A. Burgmeier,³⁷ A. Cakir,³⁷ L. Calligaris,³⁷ A. Campbell,³⁷ S. Choudhury,³⁷ F. Costanza,³⁷ C. Diez Pardos,³⁷ S. Dooling,³⁷ T. Dorland,³⁷ G. Eckerlin,³⁷ D. Eckstein,³⁷ T. Eichhorn,³⁷ G. Flucke,³⁷ J. Garay Garcia,³⁷ A. Geiser,³⁷ P. Gunnellini,³⁷ J. Hauk,³⁷ G. Hellwig,³⁷ M. Hempel,³⁷ D. Horton,³⁷ H. Jung,³⁷ M. Kasemann,³⁷ P. Katsas,³⁷ J. Kieseler,³⁷ C. Kleinwort,³⁷ D. Krücker,³⁷ W. Lange,³⁷ J. Leonard,³⁷ K. Lipka,³⁷ A. Lobanov,³⁷ W. Lohmann,^{37,p} B. Lutz,³⁷ R. Mankel,³⁷ I. Marfin,³⁷ I.-A. Melzer-Pellmann,³⁷ A. B. Meyer,³⁷ J. Mnich,³⁷ A. Mussgiller,³⁷ S. Naumann-Emme,³⁷ O. Novgorodova,³⁷ F. Nowak,³⁷ E. Ntomari,³⁷ H. Perrey,³⁷ D. Pitzl,³⁷ R. Placakyte,³⁷ A. Raspereza,³⁷ P. M. Ribeiro Cipriano,³⁷ E. Ron,³⁷ M. O. Sahin,³⁷ J. Salfeld-Nebgen,³⁷ P. Saxena,³⁷ R. Schmidt,^{37,p} T. Schoerner-Sadenius,³⁷ M. Schröder,³⁷ S. Spannagel,³⁷ A. D. R. Vargas Trevino,³⁷ R. Walsh,³⁷ C. Wissing,³⁷ M. Aldaya Martin,³⁸ V. Blobel,³⁸ M. Centis Vignali,³⁸ J. Erfle,³⁸ E. Garutti,³⁸ K. Goebel,³⁸ M. Görner,³⁸ M. Gosselink,³⁸ J. Haller,³⁸ R. S. Höing,³⁸ H. Kirschenmann,³⁸ R. Klanner,³⁸ R. Kogler,³⁸ J. Lange,³⁸ T. Lapsien,³⁸ T. Lenz,³⁸ I. Marchesini,³⁸ J. Ott,³⁸ T. Peiffer,³⁸ N. Pietsch,³⁸ D. Rathjens,³⁸ C. Sander,³⁸ H. Schettler,³⁸ P. Schleper,³⁸ E. Schlieckau,³⁸ A. Schmidt,³⁸ M. Seidel,³⁸ J. Sibille,^{38,q} V. Sola,³⁸ H. Stadie,³⁸ G. Steinbrück,³⁸ D. Troendle,³⁸ E. Usai,³⁸ L. Vanelderen,³⁸ C. Barth,³⁹ C. Baus,³⁹ J. Berger,³⁹ C. Böser,³⁹ E. Butz,³⁹ T. Chwalek,³⁹ W. De Boer,³⁹ A. Descroix,³⁹ A. Dierlamm,³⁹ M. Feindt,³⁹ F. Hartmann,^{39,c} T. Hauth,^{39,c} U. Husemann,³⁹ I. Katkov,^{39,f} A. Kornmayer,^{39,c} E. Kuznetsova,³⁹ P. Lobelle Pardo,³⁹ M. U. Mozer,³⁹ T. Müller,³⁹ A. Nürnberg,³⁹ G. Quast,³⁹ K. Rabbertz,³⁹ F. Ratnikov,³⁹ S. Röcker,³⁹ H. J. Simonis,³⁹ F. M. Stober,³⁹ R. Ulrich,³⁹ J. Wagner-Kuhr,³⁹ S. Wayand,³⁹ T. Weiler,³⁹ R. Wolf,³⁹ G. Anagnostou,⁴⁰ G. Daskalakis,⁴⁰ T. Gerasis,⁴⁰ V. A. Giakoumopoulou,⁴⁰ A. Kyriakis,⁴⁰ D. Loukas,⁴⁰ A. Markou,⁴⁰ C. Markou,⁴⁰ A. Psallidas,⁴⁰ I. Topsis-Giotis,⁴⁰ L. Gouskos,⁴¹ A. Panagiotou,⁴¹ N. Saoulidou,⁴¹ E. Stiliaris,⁴¹ X. Aslanoglou,⁴² I. Evangelou,⁴² G. Flouris,⁴² C. Foudas,⁴² P. Kokkas,⁴² N. Manthos,⁴² I. Papadopoulos,⁴² E. Paradas,⁴² G. Bencze,⁴³ C. Hajdu,⁴³ P. Hidas,⁴³ D. Horvath,^{43,r} F. Sikler,⁴³ V. Veszpremi,⁴³ G. Vesztergombi,^{43,s} A. J. Zsigmond,⁴³ N. Beni,⁴⁴ S. Czellar,⁴⁴ J. Karancsi,^{44,t} J. Molnar,⁴⁴ J. Palinkas,⁴⁴ Z. Szillasi,⁴⁴ P. Raics,⁴⁵ Z. L. Trocsanyi,⁴⁵ B. Ujvari,⁴⁵ S. K. Swain,⁴⁶ S. B. Beri,⁴⁷ V. Bhatnagar,⁴⁷ N. Dhingra,⁴⁷ R. Gupta,⁴⁷ A. K. Kalsi,⁴⁷ M. Kaur,⁴⁷ M. Mittal,⁴⁷ N. Nishu,⁴⁷ J. B. Singh,⁴⁷ A. Kumar,⁴⁸ A. Kumar,⁴⁸ S. Ahuja,⁴⁸ A. Bhardwaj,⁴⁸ B. C. Choudhary,⁴⁸ A. Kumar,⁴⁸ S. Malhotra,⁴⁸ M. Naimuddin,⁴⁸ K. Ranjan,⁴⁸ V. Sharma,⁴⁸ S. Banerjee,⁴⁹ S. Bhattacharya,⁴⁹ K. Chatterjee,⁴⁹ S. Dutta,⁴⁹ B. Gumber,⁴⁹ S. Jain,⁴⁹ S. Jain,⁴⁹ R. Khurana,⁴⁹ A. Modak,⁴⁹ S. Mukherjee,⁴⁹ D. Roy,⁴⁹ S. Sarkar,⁴⁹ M. Sharan,⁴⁹ A. Abdulsalam,⁵⁰ D. Dutta,⁵⁰ S. Kailas,⁵⁰ V. Kumar,⁵⁰ A. K. Mohanty,^{50,c} L. M. Pant,⁵⁰ P. Shukla,⁵⁰ A. Topkar,⁵⁰ T. Aziz,⁵¹ S. Banerjee,⁵¹ R. M. Chatterjee,⁵¹ R. K. Dewanjee,⁵¹ S. Dugad,⁵¹ S. Ganguly,⁵¹ S. Ghosh,⁵¹ M. Guchait,⁵¹ A. Gurtu,^{51,u} G. Kole,⁵¹ S. Kumar,⁵¹ M. Maity,^{51,v} G. Majumder,⁵¹ K. Mazumdar,⁵¹ G. B. Mohanty,⁵¹ B. Parida,⁵¹ K. Sudhakar,⁵¹ N. Wickramage,^{51,w} H. Bakhshiansohi,⁵² H. Behnamian,⁵² S. M. Etesami,^{52,x} A. Fahim,^{52,y} R. Goldouzian,⁵² A. Jafari,⁵² M. Khakzad,⁵² M. Mohammadi Najafabadi,⁵² M. Naseri,⁵² S. Paktinat Mehdiabadi,⁵² B. Safarzadeh,^{52,z} M. Zeinali,⁵² M. Felcini,⁵³ M. Grunewald,⁵³ M. Abbrescia,^{54a,54b} L. Barbone,^{54a,54b} C. Calabria,^{54a,54b} S. S. Chhibra,^{54a,54b} A. Colaleo,^{54a} D. Creanza,^{54a,54c} N. De Filippis,^{54a,54c} M. De Palma,^{54a,54b} L. Fiore,^{54a} G. Iaselli,^{54a,54c} G. Maggi,^{54a,54c} M. Maggi,^{54a} S. My,^{54a,54c} S. Nuzzo,^{54a,54b} A. Pompili,^{54a,54b}

G. Pugliese,^{54a,54c} R. Radogna,^{54a,54b,c} G. Selvaggi,^{54a,54b} L. Silvestris,^{54a,c} G. Singh,^{54a,54b} R. Venditti,^{54a,54b} P. Verwilligen,^{54a}
 G. Zito,^{54a} G. Abbiendi,^{55a} A. C. Benvenuti,^{55a} D. Bonacorsi,^{55a,55b} S. Braibant-Giacomelli,^{55a,55b} L. Brigliadori,^{55a,55b}
 R. Campanini,^{55a,55b} P. Capiluppi,^{55a,55b} A. Castro,^{55a,55b} F. R. Cavallo,^{55a} G. Codispoti,^{55a,55b} M. Cuffiani,^{55a,55b}
 G. M. Dallavalle,^{55a} F. Fabbri,^{55a} A. Fanfani,^{55a,55b} D. Fasanella,^{55a,55b} P. Giacomelli,^{55a} C. Grandi,^{55a} L. Guiducci,^{55a,55b}
 S. Marcellini,^{55a} G. Masetti,^{55a,c} A. Montanari,^{55a} F. L. Navarria,^{55a,55b} A. Perrotta,^{55a} F. Primavera,^{55a,55b} A. M. Rossi,^{55a,55b}
 T. Rovelli,^{55a,55b} G. P. Siroli,^{55a,55b} N. Tosi,^{55a,55b} R. Travaglini,^{55a,55b} S. Albergo,^{56a,56b} G. Cappello,^{56a} M. Chiorboli,^{56a,56b}
 S. Costa,^{56a,56b} F. Giordano,^{56a,56c,c} R. Potenza,^{56a,56b} A. Tricomi,^{56a,56b} C. Tuve,^{56a,56b} G. Barbagli,^{57a} V. Ciulli,^{57a,57b}
 C. Civinini,^{57a} R. D'Alessandro,^{57a,57b} E. Focardi,^{57a,57b} E. Gallo,^{57a} S. Gonzi,^{57a,57b} V. Gori,^{57a,57b,c} P. Lenzi,^{57a,57b}
 M. Meschini,^{57a} S. Paoletti,^{57a} G. Sguazzoni,^{57a} A. Tropiano,^{57a,57b} L. Benussi,⁵⁸ S. Bianco,⁵⁸ F. Fabbri,⁵⁸ D. Piccolo,⁵⁸
 F. Ferro,^{59a} M. Lo Vetere,^{59a,59b} E. Robutti,^{59a} S. Tosi,^{59a,59b} M. E. Dinardo,^{60a,60b} S. Fiorendi,^{60a,60b,c} S. Gennai,^{60a,c}
 R. Gerosa,^{60a,c} A. Ghezzi,^{60a,60b} P. Govoni,^{60a,60b} M. T. Lucchini,^{60a,60b,c} S. Malvezzi,^{60a} R. A. Manzoni,^{60a,60b}
 A. Martelli,^{60a,60b} B. Marzocchi,^{60a} D. Menasce,^{60a} L. Moroni,^{60a} M. Paganoni,^{60a,60b} D. Pedrini,^{60a} S. Ragazzi,^{60a,60b}
 N. Redaelli,^{60a} T. Tabarelli de Fatis,^{60a,60b} S. Buontempo,^{61a} N. Cavallo,^{61a,61c} S. Di Guida,^{61a,61d,c} F. Fabozzi,^{61a,61c}
 A. O. M. Iorio,^{61a,61b} L. Lista,^{61a} S. Meola,^{61a,61d,c} M. Merola,^{61a} P. Paolucci,^{61a,c} P. Azzi,^{62a} N. Bacchetta,^{62a} D. Bisello,^{62a,62b}
 A. Branca,^{62a,62b} R. Carlin,^{62a,62b} M. Dall'Osso,^{62a,62b} T. Dorigo,^{62a} M. Galanti,^{62a,62b} F. Gasparini,^{62a,62b} P. Giubilateo,^{62a,62b}
 A. Gozzelino,^{62a} K. Kanishchev,^{62a,62c} S. Lacaprarà,^{62a} M. Margoni,^{62a,62b} A. T. Meneguzzo,^{62a,62b} F. Montecassiano,^{62a}
 M. Passaseo,^{62a} J. Pazzini,^{62a,62b} N. Pozzobon,^{62a,62b} P. Ronchese,^{62a,62b} F. Simonetto,^{62a,62b} E. Torassa,^{62a} M. Tosi,^{62a,62b}
 S. Vanini,^{62a,62b} P. Zotto,^{62a,62b} A. Zucchetta,^{62a,62b} G. Zumerle,^{62a,62b} M. Gabusi,^{63a,63b} S. P. Ratti,^{63a,63b} C. Riccardi,^{63a,63b}
 P. Salvini,^{63a} P. Vitulo,^{63a,63b} M. Biasini,^{64a,64b} G. M. Bilei,^{64a} D. Ciangottini,^{64a,64b} L. Fanò,^{64a,64b} P. Lariccia,^{64a,64b}
 G. Mantovani,^{64a,64b} M. Menichelli,^{64a} F. Romeo,^{64a,64b} A. Saha,^{64a} A. Santocchia,^{64a,64b} A. Spiezia,^{64a,64b,c} K. Androsov,^{65a,aa}
 P. Azzurri,^{65a} G. Bagliesi,^{65a} J. Bernardini,^{65a} T. Boccali,^{65a} G. Broccolo,^{65a,65c} R. Castaldi,^{65a} M. A. Ciocci,^{65a,aa}
 R. Dell'Orso,^{65a} S. Donato,^{65a,65c} F. Fiori,^{65a,65c} L. Foà,^{65a,65c} A. Giassi,^{65a} M. T. Grippo,^{65a,aa} F. Ligabue,^{65a,65c}
 T. Lomtadze,^{65a} L. Martini,^{65a,65b} A. Messineo,^{65a,65b} C. S. Moon,^{65a,bb} F. Palla,^{65a,c} A. Rizzi,^{65a,65b} A. Savoy-Navarro,^{65a,cc}
 A. T. Serban,^{65a} P. Spagnolo,^{65a} P. Squillacioti,^{65a,aa} R. Tenchini,^{65a} G. Tonelli,^{65a,65b} A. Venturi,^{65a} P. G. Verdini,^{65a}
 C. Vernieri,^{65a,65c,c} L. Barone,^{66a,66b} F. Cavallari,^{66a} D. Del Re,^{66a,66b} M. Diemoz,^{66a} M. Grassi,^{66a,66b} C. Jorda,^{66a}
 E. Longo,^{66a,66b} F. Margaroli,^{66a,66b} P. Meridiani,^{66a} F. Micheli,^{66a,66b,c} S. Nourbakhsh,^{66a,66b} G. Organtini,^{66a,66b}
 R. Paramatti,^{66a} S. Rahatlou,^{66a,66b} C. Rovelli,^{66a} F. Santanastasio,^{66a,66b} L. Soffi,^{66a,66b,c} P. Traczyk,^{66a,66b} N. Amapane,^{67a,67b}
 R. Arcidiacono,^{67a,67c} S. Argiro,^{67a,67b,c} M. Arneodo,^{67a,67c} R. Bellan,^{67a,67b} C. Biino,^{67a} N. Cartiglia,^{67a} S. Casasso,^{67a,67b,c}
 M. Costa,^{67a,67b} A. Degano,^{67a,67b} N. Demaria,^{67a} L. Finco,^{67a,67b} C. Mariotti,^{67a} S. Maselli,^{67a} E. Migliore,^{67a,67b}
 V. Monaco,^{67a,67b} M. Musich,^{67a} M. M. Obertino,^{67a,67c,c} G. Ortona,^{67a,67b} L. Pacher,^{67a,67b} N. Pastrone,^{67a} M. Pelliccioni,^{67a}
 G. L. Pinna Angioni,^{67a,67b} A. Potenza,^{67a,67b} A. Romero,^{67a,67b} M. Ruspa,^{67a,67c} R. Sacchi,^{67a,67b} A. Solano,^{67a,67b}
 A. Staiano,^{67a} U. Tamponi,^{67a} S. Belforte,^{68a} V. Candelise,^{68a,68b} M. Casarsa,^{68a} F. Cossutti,^{68a} G. Della Ricca,^{68a,68b}
 B. Gobbo,^{68a} C. La Licata,^{68a,68b} M. Marone,^{68a,68b} D. Montanino,^{68a,68b} A. Schizzi,^{68a,68b,c} T. Umer,^{68a,68b} A. Zanetti,^{68a}
 S. Chang,⁶⁹ A. Kropivnitskaya,⁶⁹ S. K. Nam,⁶⁹ D. H. Kim,⁷⁰ G. N. Kim,⁷⁰ M. S. Kim,⁷⁰ D. J. Kong,⁷⁰ S. Lee,⁷⁰ Y. D. Oh,⁷⁰
 H. Park,⁷⁰ A. Sakharov,⁷⁰ D. C. Son,⁷⁰ J. Y. Kim,⁷¹ S. Song,⁷¹ S. Choi,⁷² D. Gyun,⁷² B. Hong,⁷² M. Jo,⁷² H. Kim,⁷² Y. Kim,⁷²
 B. Lee,⁷² K. S. Lee,⁷² S. K. Park,⁷² Y. Roh,⁷² M. Choi,⁷³ J. H. Kim,⁷³ I. C. Park,⁷³ S. Park,⁷³ G. Ryu,⁷³ M. S. Ryu,⁷³
 Y. Choi,⁷⁴ Y. K. Choi,⁷⁴ J. Goh,⁷⁴ E. Kwon,⁷⁴ J. Lee,⁷⁴ H. Seo,⁷⁴ I. Yu,⁷⁴ A. Juodagalvis,⁷⁵ J. R. Komaragiri,⁷⁶
 H. Castilla-Valdez,⁷⁷ E. De La Cruz-Burelo,⁷⁷ I. Heredia-de La Cruz,^{77,dd} R. Lopez-Fernandez,⁷⁷ A. Sanchez-Hernandez,⁷⁷
 S. Carrillo Moreno,⁷⁸ F. Vazquez Valencia,⁷⁸ I. Pedraza,⁷⁹ H. A. Salazar Ibarquen,⁷⁹ E. Casimiro Linares,⁸⁰
 A. Morelos Pineda,⁸⁰ D. Krofcheck,⁸¹ P. H. Butler,⁸² S. Reucroft,⁸² A. Ahmad,⁸³ M. Ahmad,⁸³ Q. Hassan,⁸³ H. R. Hoorani,⁸³
 S. Khalid,⁸³ W. A. Khan,⁸³ T. Khurshid,⁸³ M. A. Shah,⁸³ M. Shoaib,⁸³ H. Bialkowska,⁸⁴ M. Bluj,⁸⁴ B. Boimska,⁸⁴
 T. Frueboes,⁸⁴ M. Górski,⁸⁴ M. Kazana,⁸⁴ K. Nawrocki,⁸⁴ K. Romanowska-Rybinska,⁸⁴ M. Szleper,⁸⁴ P. Zalewski,⁸⁴
 G. Brona,⁸⁵ K. Bunkowski,⁸⁵ M. Cwiok,⁸⁵ W. Dominik,⁸⁵ K. Doroba,⁸⁵ A. Kalinowski,⁸⁵ M. Konecki,⁸⁵ J. Krolikowski,⁸⁵
 M. Misiura,⁸⁵ M. Olszewski,⁸⁵ W. Wolszczak,⁸⁵ P. Bargassa,⁸⁶ C. Beirão Da Cruz E Silva,⁸⁶ P. Faccioli,⁸⁶
 P. G. Ferreira Parracho,⁸⁶ M. Gallinaro,⁸⁶ F. Nguyen,⁸⁶ J. Rodrigues Antunes,⁸⁶ J. Seixas,⁸⁶ J. Varela,⁸⁶ P. Vischia,⁸⁶
 S. Afanasiev,⁸⁷ P. Bunin,⁸⁷ M. Gavrilenko,⁸⁷ I. Golutvin,⁸⁷ V. Karjavin,⁸⁷ V. Konoplyanikov,⁸⁷ A. Lanev,⁸⁷ A. Malakhov,⁸⁷
 V. Matveev,^{87,ee} P. Moiseenz,⁸⁷ V. Palichik,⁸⁷ V. Perelygin,⁸⁷ M. Savina,⁸⁷ S. Shmatov,⁸⁷ S. Shulha,⁸⁷ N. Skatchkov,⁸⁷
 V. Smirnov,⁸⁷ A. Zarubin,⁸⁷ V. Golovtsov,⁸⁸ Y. Ivanov,⁸⁸ V. Kim,^{88,ff} P. Levchenko,⁸⁸ V. Murzin,⁸⁸ V. Oreshkin,⁸⁸
 I. Smirnov,⁸⁸ V. Sulimov,⁸⁸ L. Uvarov,⁸⁸ S. Vavilov,⁸⁸ A. Vorobyev,⁸⁸ A. Vorobyev,⁸⁸ Y. Andreev,⁸⁹ A. Dermenev,⁸⁹

S. Gninenko,⁸⁹ N. Golubev,⁸⁹ M. Kirsanov,⁸⁹ N. Krasnikov,⁸⁹ A. Pashenkov,⁸⁹ D. Tlisov,⁸⁹ A. Toropin,⁸⁹ V. Epshteyn,⁹⁰ V. Gavrilov,⁹⁰ N. Lychkovskaya,⁹⁰ V. Popov,⁹⁰ G. Safronov,⁹⁰ S. Semenov,⁹⁰ A. Spiridonov,⁹⁰ V. Stolin,⁹⁰ E. Vlasov,⁹⁰ A. Zhokin,⁹⁰ V. Andreev,⁹¹ M. Azarkin,⁹¹ I. Dremin,⁹¹ M. Kirakosyan,⁹¹ A. Leonidov,⁹¹ G. Mesyats,⁹¹ S. V. Rusakov,⁹¹ A. Vinogradov,⁹¹ A. Belyaev,⁹² E. Boos,⁹² M. Dubinin,^{92,h} L. Dudko,⁹² A. Ershov,⁹² A. Gribushin,⁹² V. Klyukhin,⁹² O. Kodolova,⁹² I. Lokhtin,⁹² S. Obraztsov,⁹² S. Petrushanko,⁹² V. Savrin,⁹² A. Snigirev,⁹² I. Azhgirey,⁹³ I. Bayshev,⁹³ S. Bitioukov,⁹³ V. Kachanov,⁹³ A. Kalinin,⁹³ D. Konstantinov,⁹³ V. Krychkin,⁹³ V. Petrov,⁹³ R. Ryutin,⁹³ A. Sobol,⁹³ L. Tourtchanovitch,⁹³ S. Troshin,⁹³ N. Tyurin,⁹³ A. Uzunian,⁹³ A. Volkov,⁹³ P. Adzic,^{94,gg} M. Dordevic,⁹⁴ M. Ekmedzic,⁹⁴ J. Milosevic,⁹⁴ J. Alcaraz Maestre,⁹⁵ C. Battilana,⁹⁵ E. Calvo,⁹⁵ M. Cerrada,⁹⁵ M. Chamizo Llatas,^{95,c} N. Colino,⁹⁵ B. De La Cruz,⁹⁵ A. Delgado Peris,⁹⁵ D. Domínguez Vázquez,⁹⁵ A. Escalante Del Valle,⁹⁵ C. Fernandez Bedoya,⁹⁵ J. P. Fernández Ramos,⁹⁵ J. Flix,⁹⁵ M. C. Fouz,⁹⁵ P. Garcia-Abia,⁹⁵ O. Gonzalez Lopez,⁹⁵ S. Goy Lopez,⁹⁵ J. M. Hernandez,⁹⁵ M. I. Josa,⁹⁵ G. Merino,⁹⁵ E. Navarro De Martino,⁹⁵ A. Pérez-Calero Yzquierdo,⁹⁵ J. Puerta Pelayo,⁹⁵ A. Quintario Olmeda,⁹⁵ I. Redondo,⁹⁵ L. Romero,⁹⁵ M. S. Soares,⁹⁵ C. Albajar,⁹⁶ J. F. de Trocóniz,⁹⁶ M. Missiroli,⁹⁶ H. Brun,⁹⁷ J. Cuevas,⁹⁷ J. Fernandez Menendez,⁹⁷ S. Folgueras,⁹⁷ I. Gonzalez Caballero,⁹⁷ L. Lloret Iglesias,⁹⁷ J. A. Brochero Cifuentes,⁹⁸ I. J. Cabrillo,⁹⁸ A. Calderon,⁹⁸ J. Duarte Campderros,⁹⁸ M. Fernandez,⁹⁸ G. Gomez,⁹⁸ A. Graziano,⁹⁸ A. Lopez Virto,⁹⁸ J. Marco,⁹⁸ R. Marco,⁹⁸ C. Martinez Rivero,⁹⁸ F. Matorras,⁹⁸ F. J. Munoz Sanchez,⁹⁸ J. Piedra Gomez,⁹⁸ T. Rodrigo,⁹⁸ A. Y. Rodríguez-Marrero,⁹⁸ A. Ruiz-Jimeno,⁹⁸ L. Scodellaro,⁹⁸ I. Vila,⁹⁸ R. Vilar Cortabitarte,⁹⁸ D. Abbaneo,⁹⁹ E. Auffray,⁹⁹ G. Auzinger,⁹⁹ M. Bachtis,⁹⁹ P. Baillon,⁹⁹ A. H. Ball,⁹⁹ D. Barney,⁹⁹ A. Benaglia,⁹⁹ J. Bendavid,⁹⁹ L. Benhabib,⁹⁹ J. F. Benitez,⁹⁹ C. Bernet,^{99,i} G. Bianchi,⁹⁹ P. Bloch,⁹⁹ A. Bocci,⁹⁹ A. Bonato,⁹⁹ O. Bondu,⁹⁹ C. Botta,⁹⁹ H. Breuker,⁹⁹ T. Camporesi,⁹⁹ G. Cerminara,⁹⁹ T. Christiansen,⁹⁹ S. Colafranceschi,^{99,hh} M. D'Alfonso,⁹⁹ D. d'Enterria,⁹⁹ A. Dabrowski,⁹⁹ A. David,⁹⁹ F. De Guio,⁹⁹ A. De Roeck,⁹⁹ S. De Visscher,⁹⁹ M. Dobson,⁹⁹ N. Dupont-Sagorin,⁹⁹ A. Elliott-Peisert,⁹⁹ J. Eugster,⁹⁹ G. Franzoni,⁹⁹ W. Funk,⁹⁹ M. Giffels,⁹⁹ D. Gigi,⁹⁹ K. Gill,⁹⁹ D. Giordano,⁹⁹ M. Girone,⁹⁹ F. Glege,⁹⁹ R. Guida,⁹⁹ S. Gundacker,⁹⁹ M. Guthoff,⁹⁹ J. Hammer,⁹⁹ M. Hansen,⁹⁹ P. Harris,⁹⁹ J. Hegeman,⁹⁹ V. Innocente,⁹⁹ P. Janot,⁹⁹ K. Kousouris,⁹⁹ K. Krajczar,⁹⁹ P. Lecoq,⁹⁹ C. Lourenço,⁹⁹ N. Magini,⁹⁹ L. Malgeri,⁹⁹ M. Mannelli,⁹⁹ L. Masetti,⁹⁹ F. Meijers,⁹⁹ S. Mersi,⁹⁹ E. Meschi,⁹⁹ F. Moortgat,⁹⁹ S. Morovic,⁹⁹ M. Mulders,⁹⁹ P. Musella,⁹⁹ L. Orsini,⁹⁹ L. Pape,⁹⁹ E. Perez,⁹⁹ L. Perrozzi,⁹⁹ A. Petrilli,⁹⁹ G. Petrucciani,⁹⁹ A. Pfeiffer,⁹⁹ M. Pierini,⁹⁹ M. Pimiä,⁹⁹ D. Piparo,⁹⁹ M. Plagge,⁹⁹ A. Racz,⁹⁹ G. Rolandi,^{99,ii} M. Rovere,⁹⁹ H. Sakulin,⁹⁹ C. Schäfer,⁹⁹ C. Schwick,⁹⁹ S. Sekmen,⁹⁹ A. Sharma,⁹⁹ P. Siegrist,⁹⁹ P. Silva,⁹⁹ M. Simon,⁹⁹ P. Sphicas,^{99,jj} D. Spiga,⁹⁹ J. Steggemann,⁹⁹ B. Stieger,⁹⁹ M. Stoye,⁹⁹ D. Treille,⁹⁹ A. Tsiros,⁹⁹ G. I. Veres,^{99,s} J. R. Vlimant,⁹⁹ N. Wardle,⁹⁹ H. K. Wöhri,⁹⁹ W. D. Zeuner,⁹⁹ W. Bertl,¹⁰⁰ K. Deiters,¹⁰⁰ W. Erdmann,¹⁰⁰ R. Horisberger,¹⁰⁰ Q. Ingram,¹⁰⁰ H. C. Kaestli,¹⁰⁰ S. König,¹⁰⁰ D. Kotlinski,¹⁰⁰ U. Langenegger,¹⁰⁰ D. Renker,¹⁰⁰ T. Rohe,¹⁰⁰ F. Bachmair,¹⁰¹ L. Bäni,¹⁰¹ L. Bianchini,¹⁰¹ P. Bortignon,¹⁰¹ M. A. Buchmann,¹⁰¹ B. Casal,¹⁰¹ N. Chanon,¹⁰¹ A. Deisher,¹⁰¹ G. Dissertori,¹⁰¹ M. Dittmar,¹⁰¹ M. Donegà,¹⁰¹ M. Dünser,¹⁰¹ P. Eller,¹⁰¹ C. Grab,¹⁰¹ D. Hits,¹⁰¹ W. Lustermann,¹⁰¹ B. Mangano,¹⁰¹ A. C. Marini,¹⁰¹ P. Martinez Ruiz del Arbol,¹⁰¹ D. Meister,¹⁰¹ N. Mohr,¹⁰¹ C. Nägeli,^{101,kk} P. Nef,¹⁰¹ F. Nessi-Tedaldi,¹⁰¹ F. Pandolfi,¹⁰¹ F. Paus,¹⁰¹ M. Peruzzi,¹⁰¹ M. Quittnat,¹⁰¹ L. Rebane,¹⁰¹ F. J. Ronga,¹⁰¹ M. Rossini,¹⁰¹ A. Starodumov,^{101,ll} M. Takahashi,¹⁰¹ K. Theofilatos,¹⁰¹ R. Wallny,¹⁰¹ H. A. Weber,¹⁰¹ C. AMSler,^{102,mmm} M. F. Canelli,¹⁰² V. Chiochia,¹⁰² A. De Cosa,¹⁰² A. Hinzmann,¹⁰² T. Hreus,¹⁰² M. Ivova Rikova,¹⁰² B. Kilminster,¹⁰² B. Millan Mejias,¹⁰² J. Ngadiuba,¹⁰² P. Robmann,¹⁰² H. Snoek,¹⁰² S. Taroni,¹⁰² M. Verzetti,¹⁰² Y. Yang,¹⁰² M. Cardaci,¹⁰³ K. H. Chen,¹⁰³ C. Ferro,¹⁰³ C. M. Kuo,¹⁰³ W. Lin,¹⁰³ Y. J. Lu,¹⁰³ R. Volpe,¹⁰³ S. S. Yu,¹⁰³ P. Chang,¹⁰⁴ Y. H. Chang,¹⁰⁴ Y. W. Chang,¹⁰⁴ Y. Chao,¹⁰⁴ K. F. Chen,¹⁰⁴ P. H. Chen,¹⁰⁴ C. Dietz,¹⁰⁴ U. Grundler,¹⁰⁴ W.-S. Hou,¹⁰⁴ K. Y. Kao,¹⁰⁴ Y. J. Lei,¹⁰⁴ Y. F. Liu,¹⁰⁴ R.-S. Lu,¹⁰⁴ D. Majumder,¹⁰⁴ E. Petrakou,¹⁰⁴ X. Shi,¹⁰⁴ Y. M. Tzeng,¹⁰⁴ R. Wilken,¹⁰⁴ B. Asavapibhop,¹⁰⁵ N. Srimanobhas,¹⁰⁵ N. Suwonjandee,¹⁰⁵ A. Adiguzel,¹⁰⁶ M. N. Bakirci,^{106,nn} S. Cerci,^{106,oo} C. Dozen,¹⁰⁶ I. Dumanoglu,¹⁰⁶ E. Eskut,¹⁰⁶ S. Girgis,¹⁰⁶ G. Gokbulut,¹⁰⁶ E. Gurpınar,¹⁰⁶ I. Hos,¹⁰⁶ E. E. Kangal,¹⁰⁶ A. Kayis Topaksu,¹⁰⁶ G. Onengut,^{106,pp} K. Ozdemir,¹⁰⁶ S. Ozturk,^{106,nn} A. Polatoz,¹⁰⁶ K. Sogut,^{106,qq} D. Sunar Cerci,^{106,oo} B. Tali,^{106,oo} H. Topakli,^{106,nn} M. Vergili,¹⁰⁶ I. V. Akin,¹⁰⁷ B. Bilin,¹⁰⁷ S. Bilmis,¹⁰⁷ H. Gamsizkan,¹⁰⁷ G. Karapınar,^{107,rr} K. Ocalan,¹⁰⁷ U. E. Surat,¹⁰⁷ M. Yalvac,¹⁰⁷ M. Zeyrek,¹⁰⁷ E. Gülmez,¹⁰⁸ B. Isildak,^{108,ss} M. Kaya,^{108,tt} O. Kaya,^{108,tt} H. Bahtiyar,^{109,uu} E. Barlas,¹⁰⁹ K. Cankocak,¹⁰⁹ F. I. Vardarli,¹⁰⁹ M. Yücel,¹⁰⁹ L. Levchuk,¹¹⁰ P. Sorokin,¹¹⁰ J. J. Brooke,¹¹¹ E. Clement,¹¹¹ D. Cussans,¹¹¹ H. Flacher,¹¹¹ R. Frazier,¹¹¹ J. Goldstein,¹¹¹ M. Grimes,¹¹¹ G. P. Heath,¹¹¹ H. F. Heath,¹¹¹ J. Jacob,¹¹¹ L. Kreczko,¹¹¹ C. Lucas,¹¹¹ Z. Meng,¹¹¹ D. M. Newbold,^{111,vv} S. Paramesvaran,¹¹¹ A. Poll,¹¹¹ S. Senkin,¹¹¹ V. J. Smith,¹¹¹ T. Williams,¹¹¹ K. W. Bell,¹¹² A. Belyaev,^{112,ww} C. Brew,¹¹² R. M. Brown,¹¹² D. J. A. Cockerill,¹¹² J. A. Coughlan,¹¹² K. Harder,¹¹² S. Harper,¹¹² E. Olaiya,¹¹² D. Petyt,¹¹²

C. H. Shepherd-Themistocleous,¹¹² A. Thea,¹¹² I. R. Tomalin,¹¹² W. J. Womersley,¹¹² S. D. Worm,¹¹² M. Baber,¹¹³ R. Bainbridge,¹¹³ O. Buchmuller,¹¹³ D. Burton,¹¹³ D. Colling,¹¹³ N. Cripps,¹¹³ M. Cutajar,¹¹³ P. Dauncey,¹¹³ G. Davies,¹¹³ M. Della Negra,¹¹³ P. Dunne,¹¹³ W. Ferguson,¹¹³ J. Fulcher,¹¹³ D. Futyan,¹¹³ A. Gilbert,¹¹³ G. Hall,¹¹³ G. Iles,¹¹³ M. Jarvis,¹¹³ G. Karapostoli,¹¹³ M. Kenzie,¹¹³ R. Lane,¹¹³ R. Lucas,^{113,uv} L. Lyons,¹¹³ A.-M. Magnan,¹¹³ S. Malik,¹¹³ J. Marrouche,¹¹³ B. Mathias,¹¹³ J. Nash,¹¹³ A. Nikitenko,^{113,ll} J. Pela,¹¹³ M. Pesaresi,¹¹³ K. Petridis,¹¹³ D. M. Raymond,¹¹³ S. Rogerson,¹¹³ A. Rose,¹¹³ C. Seez,¹¹³ P. Sharp,^{113,a} A. Tapper,¹¹³ M. Vazquez Acosta,¹¹³ T. Virdee,¹¹³ J. E. Cole,¹¹⁴ P. R. Hobson,¹¹⁴ A. Khan,¹¹⁴ P. Kyberd,¹¹⁴ D. Leggat,¹¹⁴ D. Leslie,¹¹⁴ W. Martin,¹¹⁴ I. D. Reid,¹¹⁴ P. Symonds,¹¹⁴ L. Teodorescu,¹¹⁴ M. Turner,¹¹⁴ J. Dittmann,¹¹⁵ K. Hatakeyama,¹¹⁵ A. Kasmi,¹¹⁵ H. Liu,¹¹⁵ T. Scarborough,¹¹⁵ O. Charaf,¹¹⁶ S. I. Cooper,¹¹⁶ C. Henderson,¹¹⁶ P. Rumerio,¹¹⁶ A. Avetisyan,¹¹⁷ T. Bose,¹¹⁷ C. Fantasia,¹¹⁷ A. Heister,¹¹⁷ P. Lawson,¹¹⁷ C. Richardson,¹¹⁷ J. Rohlf,¹¹⁷ D. Sperka,¹¹⁷ J. St. John,¹¹⁷ L. Sulak,¹¹⁷ J. Alimena,¹¹⁸ S. Bhattacharya,¹¹⁸ G. Christopher,¹¹⁸ D. Cutts,¹¹⁸ Z. Demiragli,¹¹⁸ A. Ferapontov,¹¹⁸ A. Garabedian,¹¹⁸ U. Heintz,¹¹⁸ S. Jabeen,¹¹⁸ G. Kukartsev,¹¹⁸ E. Laird,¹¹⁸ G. Landsberg,¹¹⁸ M. Luk,¹¹⁸ M. Narain,¹¹⁸ M. Segala,¹¹⁸ T. Sinthuprasith,¹¹⁸ T. Speer,¹¹⁸ J. Swanson,¹¹⁸ R. Breedon,¹¹⁹ G. Breto,¹¹⁹ M. Calderon De La Barca Sanchez,¹¹⁹ S. Chauhan,¹¹⁹ M. Chertok,¹¹⁹ J. Conway,¹¹⁹ R. Conway,¹¹⁹ P. T. Cox,¹¹⁹ R. Erbacher,¹¹⁹ M. Gardner,¹¹⁹ W. Ko,¹¹⁹ R. Lander,¹¹⁹ T. Miceli,¹¹⁹ M. Mulhearn,¹¹⁹ D. Pellett,¹¹⁹ J. Pilot,¹¹⁹ F. Ricci-Tam,¹¹⁹ M. Searle,¹¹⁹ S. Shalhout,¹¹⁹ J. Smith,¹¹⁹ M. Squires,¹¹⁹ D. Stolp,¹¹⁹ M. Tripathi,¹¹⁹ S. Wilbur,¹¹⁹ R. Yohay,¹¹⁹ R. Cousins,¹²⁰ P. Everaerts,¹²⁰ C. Farrell,¹²⁰ J. Hauser,¹²⁰ M. Ignatenko,¹²⁰ G. Rakness,¹²⁰ E. Takasugi,¹²⁰ V. Valuev,¹²⁰ M. Weber,¹²⁰ J. Babb,¹²¹ R. Clare,¹²¹ J. Ellison,¹²¹ J. W. Gary,¹²¹ G. Hanson,¹²¹ J. Heilman,¹²¹ P. Jandir,¹²¹ E. Kennedy,¹²¹ F. Lacroix,¹²¹ H. Liu,¹²¹ O. R. Long,¹²¹ A. Luthra,¹²¹ M. Malberti,¹²¹ H. Nguyen,¹²¹ A. Shrinivas,¹²¹ J. Sturdy,¹²¹ S. Sumowidagdo,¹²¹ S. Wimpenny,¹²¹ W. Andrews,¹²² J. G. Branson,¹²² G. B. Cerati,¹²² S. Cittolin,¹²² R. T. D'Agnolo,¹²² D. Evans,¹²² A. Holzner,¹²² R. Kelley,¹²² M. Lebourgeois,¹²² J. Letts,¹²² I. Macneill,¹²² D. Olivito,¹²² S. Padhi,¹²² C. Palmer,¹²² M. Pieri,¹²² M. Sani,¹²² V. Sharma,¹²² S. Simon,¹²² E. Sudano,¹²² M. Tadel,¹²² Y. Tu,¹²² A. Vartak,¹²² F. Würthwein,¹²² A. Yagil,¹²² J. Yoo,¹²² D. Barge,¹²³ J. Bradmiller-Feld,¹²³ C. Campagnari,¹²³ T. Danielson,¹²³ A. Dishaw,¹²³ K. Flowers,¹²³ M. Franco Sevilla,¹²³ P. Geffert,¹²³ C. George,¹²³ F. Golf,¹²³ J. Incandela,¹²³ C. Justus,¹²³ N. Mccoll,¹²³ J. Richman,¹²³ D. Stuart,¹²³ W. To,¹²³ C. West,¹²³ A. Apresyan,¹²⁴ A. Bornheim,¹²⁴ J. Bunn,¹²⁴ Y. Chen,¹²⁴ E. Di Marco,¹²⁴ J. Duarte,¹²⁴ A. Mott,¹²⁴ H. B. Newman,¹²⁴ C. Pena,¹²⁴ C. Rogan,¹²⁴ M. Spiropulu,¹²⁴ V. Timciuc,¹²⁴ R. Wilkinson,¹²⁴ S. Xie,¹²⁴ R. Y. Zhu,¹²⁴ V. Azzolini,¹²⁵ A. Calamba,¹²⁵ R. Carroll,¹²⁵ T. Ferguson,¹²⁵ Y. Iiyama,¹²⁵ M. Paulini,¹²⁵ J. Russ,¹²⁵ H. Vogel,¹²⁵ I. Vorobiev,¹²⁵ J. P. Cumalat,¹²⁶ B. R. Drell,¹²⁶ W. T. Ford,¹²⁶ A. Gaz,¹²⁶ E. Luigi Lopez,¹²⁶ U. Nauenberg,¹²⁶ J. G. Smith,¹²⁶ K. Stenson,¹²⁶ K. A. Ulmer,¹²⁶ S. R. Wagner,¹²⁶ J. Alexander,¹²⁷ A. Chatterjee,¹²⁷ J. Chu,¹²⁷ S. Dittmer,¹²⁷ N. Eggert,¹²⁷ W. Hopkins,¹²⁷ B. Kreis,¹²⁷ N. Mirman,¹²⁷ G. Nicolas Kaufman,¹²⁷ J. R. Patterson,¹²⁷ A. Ryd,¹²⁷ E. Salvati,¹²⁷ L. Skinnari,¹²⁷ W. Sun,¹²⁷ W. D. Teo,¹²⁷ J. Thom,¹²⁷ J. Thompson,¹²⁷ J. Tucker,¹²⁷ Y. Weng,¹²⁷ L. Winstrom,¹²⁷ P. Wittich,¹²⁷ D. Winn,¹²⁸ S. Abdullin,¹²⁹ M. Albrow,¹²⁹ J. Anderson,¹²⁹ G. Apollinari,¹²⁹ L. A. T. Bauerdick,¹²⁹ A. Beretvas,¹²⁹ J. Berryhill,¹²⁹ P. C. Bhat,¹²⁹ K. Burkett,¹²⁹ J. N. Butler,¹²⁹ H. W. K. Cheung,¹²⁹ F. Chlebana,¹²⁹ S. Cihangir,¹²⁹ V. D. Elvira,¹²⁹ I. Fisk,¹²⁹ J. Freeman,¹²⁹ E. Gottschalk,¹²⁹ L. Gray,¹²⁹ D. Green,¹²⁹ S. Grünendahl,¹²⁹ O. Gutsche,¹²⁹ J. Hanlon,¹²⁹ D. Hare,¹²⁹ R. M. Harris,¹²⁹ J. Hirschauer,¹²⁹ B. Hooberman,¹²⁹ S. Jindariani,¹²⁹ M. Johnson,¹²⁹ U. Joshi,¹²⁹ K. Kaadze,¹²⁹ B. Klima,¹²⁹ S. Kwan,¹²⁹ J. Linacre,¹²⁹ D. Lincoln,¹²⁹ R. Lipton,¹²⁹ T. Liu,¹²⁹ J. Lykken,¹²⁹ K. Maeshima,¹²⁹ J. M. Marraffino,¹²⁹ V. I. Martinez Outschoorn,¹²⁹ S. Maruyama,¹²⁹ D. Mason,¹²⁹ P. McBride,¹²⁹ K. Mishra,¹²⁹ S. Mrenna,¹²⁹ Y. Musienko,^{129,ee} S. Nahn,¹²⁹ C. Newman-Holmes,¹²⁹ V. O'Dell,¹²⁹ O. Prokofyev,¹²⁹ E. Sexton-Kennedy,¹²⁹ S. Sharma,¹²⁹ A. Soha,¹²⁹ W. J. Spalding,¹²⁹ L. Spiegel,¹²⁹ L. Taylor,¹²⁹ S. Tkaczyk,¹²⁹ N. V. Tran,¹²⁹ L. Uplegger,¹²⁹ E. W. Vaandering,¹²⁹ R. Vidal,¹²⁹ A. Whitbeck,¹²⁹ J. Whitmore,¹²⁹ F. Yang,¹²⁹ D. Acosta,¹³⁰ P. Avery,¹³⁰ D. Bourilkov,¹³⁰ M. Carver,¹³⁰ T. Cheng,¹³⁰ D. Curry,¹³⁰ S. Das,¹³⁰ M. De Gruttola,¹³⁰ G. P. Di Giovanni,¹³⁰ R. D. Field,¹³⁰ M. Fisher,¹³⁰ I. K. Furic,¹³⁰ J. Hugon,¹³⁰ J. Konigsberg,¹³⁰ A. Korytov,¹³⁰ T. Kypreos,¹³⁰ J. F. Low,¹³⁰ K. Matchev,¹³⁰ P. Milenovic,^{130,xx} G. Mitselmakher,¹³⁰ L. Muniz,¹³⁰ A. Rinkevicius,¹³⁰ L. Shchutska,¹³⁰ N. Skhirtladze,¹³⁰ M. Snowball,¹³⁰ J. Yelton,¹³⁰ M. Zakaria,¹³⁰ V. Gaultney,¹³¹ S. Hewamanage,¹³¹ S. Linn,¹³¹ P. Markowitz,¹³¹ G. Martinez,¹³¹ J. L. Rodriguez,¹³¹ T. Adams,¹³² A. Askew,¹³² J. Bochenek,¹³² B. Diamond,¹³² J. Haas,¹³² S. Hagopian,¹³² V. Hagopian,¹³² K. F. Johnson,¹³² H. Prosper,¹³² V. Veeraraghavan,¹³² M. Weinberg,¹³² M. M. Baarmand,¹³³ M. Hohmann,¹³³ H. Kalakhety,¹³³ F. Yumiceva,¹³³ M. R. Adams,¹³⁴ L. Apanasevich,¹³⁴ V. E. Bazterra,¹³⁴ D. Berry,¹³⁴ R. R. Betts,¹³⁴ I. Bucinskaite,¹³⁴ R. Cavanaugh,¹³⁴ O. Evdokimov,¹³⁴ L. Gauthier,¹³⁴ C. E. Gerber,¹³⁴ D. J. Hofman,¹³⁴ S. Khalatyan,¹³⁴ P. Kurt,¹³⁴ D. H. Moon,¹³⁴ C. O'Brien,¹³⁴ C. Silkworth,¹³⁴ P. Turner,¹³⁴ N. Varelas,¹³⁴ E. A. Albayrak,^{135,uu} B. Bilki,^{135,yy} W. Clarida,¹³⁵ K. Dilsiz,¹³⁵ F. Duru,¹³⁵ M. Haytmyradov,¹³⁵ J.-P. Merlo,¹³⁵ H. Mermerkaya,^{135,zz}

A. Mestvirishvili,¹³⁵ A. Moeller,¹³⁵ J. Nachtman,¹³⁵ H. Ogul,¹³⁵ Y. Onel,¹³⁵ F. Ozok,^{135,uu} A. Penzo,¹³⁵ R. Rahmat,¹³⁵ S. Sen,¹³⁵ P. Tan,¹³⁵ E. Tiras,¹³⁵ J. Wetzel,¹³⁵ T. Yetkin,^{135,aaa} K. Yi,¹³⁵ B. A. Barnett,¹³⁶ B. Blumenfeld,¹³⁶ S. Bolognesi,¹³⁶ D. Fehling,¹³⁶ A. V. Gritsan,¹³⁶ P. Maksimovic,¹³⁶ C. Martin,¹³⁶ M. Swartz,¹³⁶ P. Baringer,¹³⁷ A. Bean,¹³⁷ G. Benelli,¹³⁷ C. Bruner,¹³⁷ J. Gray,¹³⁷ R. P. Kenny III,¹³⁷ M. Murray,¹³⁷ D. Noonan,¹³⁷ S. Sanders,¹³⁷ J. Sekaric,¹³⁷ R. Stringer,¹³⁷ Q. Wang,¹³⁷ J. S. Wood,¹³⁷ A. F. Barfuss,¹³⁸ I. Chakaberia,¹³⁸ A. Ivanov,¹³⁸ S. Khalil,¹³⁸ M. Makouski,¹³⁸ Y. Maravin,¹³⁸ L. K. Saini,¹³⁸ S. Shrestha,¹³⁸ I. Svintradze,¹³⁸ J. Gronberg,¹³⁹ D. Lange,¹³⁹ F. Rebassoo,¹³⁹ D. Wright,¹³⁹ A. Baden,¹⁴⁰ B. Calvert,¹⁴⁰ S. C. Eno,¹⁴⁰ J. A. Gomez,¹⁴⁰ N. J. Hadley,¹⁴⁰ R. G. Kellogg,¹⁴⁰ T. Kolberg,¹⁴⁰ Y. Lu,¹⁴⁰ M. Marionneau,¹⁴⁰ A. C. Mignerey,¹⁴⁰ K. Pedro,¹⁴⁰ A. Skuja,¹⁴⁰ M. B. Tonjes,¹⁴⁰ S. C. Tonwar,¹⁴⁰ A. Apyan,¹⁴¹ R. Barbieri,¹⁴¹ G. Bauer,¹⁴¹ W. Busza,¹⁴¹ I. A. Cali,¹⁴¹ M. Chan,¹⁴¹ L. Di Matteo,¹⁴¹ V. Dutta,¹⁴¹ G. Gomez Ceballos,¹⁴¹ M. Goncharov,¹⁴¹ D. Gulhan,¹⁴¹ M. Klute,¹⁴¹ Y. S. Lai,¹⁴¹ Y.-J. Lee,¹⁴¹ A. Levin,¹⁴¹ P. D. Luckey,¹⁴¹ T. Ma,¹⁴¹ C. Paus,¹⁴¹ D. Ralph,¹⁴¹ C. Roland,¹⁴¹ G. Roland,¹⁴¹ G. S. F. Stephans,¹⁴¹ F. Stöckli,¹⁴¹ K. Sumorok,¹⁴¹ D. Velicanu,¹⁴¹ J. Veverka,¹⁴¹ B. Wyslouch,¹⁴¹ M. Yang,¹⁴¹ M. Zanetti,¹⁴¹ V. Zhukova,¹⁴¹ B. Dahmes,¹⁴² A. De Benedetti,¹⁴² A. Gude,¹⁴² S. C. Kao,¹⁴² K. Klapoetke,¹⁴² Y. Kubota,¹⁴² J. Mans,¹⁴² N. Pastika,¹⁴² R. Rusack,¹⁴² A. Singovsky,¹⁴² N. Tambe,¹⁴² J. Turkewitz,¹⁴² J. G. Acosta,¹⁴³ S. Oliveros,¹⁴³ E. Avdeeva,¹⁴⁴ K. Bloom,¹⁴⁴ S. Bose,¹⁴⁴ D. R. Claes,¹⁴⁴ A. Dominguez,¹⁴⁴ R. Gonzalez Suarez,¹⁴⁴ J. Keller,¹⁴⁴ D. Knowlton,¹⁴⁴ I. Kravchenko,¹⁴⁴ J. Lazo-Flores,¹⁴⁴ S. Malik,¹⁴⁴ F. Meier,¹⁴⁴ G. R. Snow,¹⁴⁴ J. Dolen,¹⁴⁵ A. Godshalk,¹⁴⁵ I. Iashvili,¹⁴⁵ A. Kharchilava,¹⁴⁵ A. Kumar,¹⁴⁵ S. Rappoccio,¹⁴⁵ G. Alverson,¹⁴⁶ E. Barberis,¹⁴⁶ D. Baumgartel,¹⁴⁶ M. Chasco,¹⁴⁶ J. Haley,¹⁴⁶ A. Massironi,¹⁴⁶ D. M. Morse,¹⁴⁶ D. Nash,¹⁴⁶ T. Orimoto,¹⁴⁶ D. Trocino,¹⁴⁶ D. Wood,¹⁴⁶ J. Zhang,¹⁴⁶ K. A. Hahn,¹⁴⁷ A. Kubik,¹⁴⁷ N. Mucia,¹⁴⁷ N. Odell,¹⁴⁷ B. Pollack,¹⁴⁷ A. Pozdnyakov,¹⁴⁷ M. Schmitt,¹⁴⁷ S. Stoynev,¹⁴⁷ K. Sung,¹⁴⁷ M. Velasco,¹⁴⁷ S. Won,¹⁴⁷ A. Brinkerhoff,¹⁴⁸ K. M. Chan,¹⁴⁸ A. Drozdetskiy,¹⁴⁸ M. Hildreth,¹⁴⁸ C. Jessop,¹⁴⁸ D. J. Karmgard,¹⁴⁸ N. Kellams,¹⁴⁸ K. Lannon,¹⁴⁸ W. Luo,¹⁴⁸ S. Lynch,¹⁴⁸ N. Marinelli,¹⁴⁸ T. Pearson,¹⁴⁸ M. Planer,¹⁴⁸ R. Ruchti,¹⁴⁸ N. Valls,¹⁴⁸ M. Wayne,¹⁴⁸ M. Wolf,¹⁴⁸ A. Woodard,¹⁴⁸ L. Antonelli,¹⁴⁹ J. Brinson,¹⁴⁹ B. Bylsma,¹⁴⁹ L. S. Durkin,¹⁴⁹ S. Flowers,¹⁴⁹ C. Hill,¹⁴⁹ R. Hughes,¹⁴⁹ K. Kotov,¹⁴⁹ T. Y. Ling,¹⁴⁹ D. Puigh,¹⁴⁹ M. Rodenburg,¹⁴⁹ G. Smith,¹⁴⁹ C. Vuosalo,¹⁴⁹ B. L. Winer,¹⁴⁹ H. Wolfe,¹⁴⁹ H. W. Wulsin,¹⁴⁹ E. Berry,¹⁵⁰ O. Driga,¹⁵⁰ P. Elmer,¹⁵⁰ P. Hebda,¹⁵⁰ A. Hunt,¹⁵⁰ S. A. Koay,¹⁵⁰ P. Lujan,¹⁵⁰ D. Marlow,¹⁵⁰ T. Medvedeva,¹⁵⁰ M. Mooney,¹⁵⁰ J. Olsen,¹⁵⁰ P. Piroué,¹⁵⁰ X. Quan,¹⁵⁰ H. Saka,¹⁵⁰ D. Stickland,^{150,c} C. Tully,¹⁵⁰ J. S. Werner,¹⁵⁰ S. C. Zenz,¹⁵⁰ A. Zuranski,¹⁵⁰ E. Brownson,¹⁵¹ H. Mendez,¹⁵¹ J. E. Ramirez Vargas,¹⁵¹ E. Alagoz,¹⁵² V. E. Barnes,¹⁵² D. Benedetti,¹⁵² G. Bolla,¹⁵² D. Bortoletto,¹⁵² M. De Mattia,¹⁵² A. Everett,¹⁵² Z. Hu,¹⁵² M. K. Jha,¹⁵² M. Jones,¹⁵² K. Jung,¹⁵² M. Kress,¹⁵² N. Leonardo,¹⁵² D. Lopes Pegna,¹⁵² V. Maroussov,¹⁵² P. Merkel,¹⁵² D. H. Miller,¹⁵² N. Neumeister,¹⁵² B. C. Radburn-Smith,¹⁵² I. Shipsey,¹⁵² D. Silvers,¹⁵² A. Svyatkovskiy,¹⁵² F. Wang,¹⁵² W. Xie,¹⁵² L. Xu,¹⁵² H. D. Yoo,¹⁵² J. Zablocki,¹⁵² Y. Zheng,¹⁵² N. Parashar,¹⁵³ J. Stupak,¹⁵³ A. Adair,¹⁵⁴ B. Akgun,¹⁵⁴ K. M. Ecklund,¹⁵⁴ F. J. M. Geurts,¹⁵⁴ W. Li,¹⁵⁴ B. Michlin,¹⁵⁴ B. P. Padley,¹⁵⁴ R. Redjimi,¹⁵⁴ J. Roberts,¹⁵⁴ J. Zabel,¹⁵⁴ B. Betchart,¹⁵⁵ A. Bodek,¹⁵⁵ R. Covarelli,¹⁵⁵ P. de Barbaro,¹⁵⁵ R. Demina,¹⁵⁵ Y. Eshaq,¹⁵⁵ T. Ferbel,¹⁵⁵ A. Garcia-Bellido,¹⁵⁵ P. Goldenzweig,¹⁵⁵ J. Han,¹⁵⁵ A. Harel,¹⁵⁵ A. Khukhunaishvili,¹⁵⁵ D. C. Miner,¹⁵⁵ G. Petrillo,¹⁵⁵ D. Vishnevskiy,¹⁵⁵ R. Ciesielski,¹⁵⁶ L. Demortier,¹⁵⁶ K. Goulianos,¹⁵⁶ G. Lungu,¹⁵⁶ C. Mesropian,¹⁵⁶ S. Arora,¹⁵⁷ A. Barker,¹⁵⁷ J. P. Chou,¹⁵⁷ C. Contreras-Campana,¹⁵⁷ E. Contreras-Campana,¹⁵⁷ D. Duggan,¹⁵⁷ D. Ferencek,¹⁵⁷ Y. Gershtein,¹⁵⁷ R. Gray,¹⁵⁷ E. Halkiadakis,¹⁵⁷ D. Hidas,¹⁵⁷ A. Lath,¹⁵⁷ S. Panwalkar,¹⁵⁷ M. Park,¹⁵⁷ R. Patel,¹⁵⁷ V. Rekovic,¹⁵⁷ S. Salur,¹⁵⁷ S. Schnetzer,¹⁵⁷ C. Seitz,¹⁵⁷ S. Somalwar,¹⁵⁷ R. Stone,¹⁵⁷ S. Thomas,¹⁵⁷ P. Thomassen,¹⁵⁷ M. Walker,¹⁵⁷ K. Rose,¹⁵⁸ S. Spanier,¹⁵⁸ A. York,¹⁵⁸ O. Bouhali,^{159,bbb} R. Eusebi,¹⁵⁹ W. Flanagan,¹⁵⁹ J. Gilmore,¹⁵⁹ T. Kamon,^{159,ccc} V. Khotilovich,¹⁵⁹ V. Krutelyov,¹⁵⁹ R. Montalvo,¹⁵⁹ I. Osipenkov,¹⁵⁹ Y. Pakhotin,¹⁵⁹ A. Perloff,¹⁵⁹ J. Roe,¹⁵⁹ A. Rose,¹⁵⁹ A. Safonov,¹⁵⁹ T. Sakuma,¹⁵⁹ I. Suarez,¹⁵⁹ A. Tatarinov,¹⁵⁹ N. Akchurin,¹⁶⁰ C. Cowden,¹⁶⁰ J. Damgov,¹⁶⁰ C. Dragoiu,¹⁶⁰ P. R. Duderø,¹⁶⁰ J. Faulkner,¹⁶⁰ K. Kovitanggoon,¹⁶⁰ S. Kunori,¹⁶⁰ S. W. Lee,¹⁶⁰ T. Libeiro,¹⁶⁰ I. Volobouev,¹⁶⁰ E. Appelt,¹⁶¹ A. G. Delannoy,¹⁶¹ S. Greene,¹⁶¹ A. Gurrola,¹⁶¹ W. Johns,¹⁶¹ C. Maguire,¹⁶¹ Y. Mao,¹⁶¹ A. Melo,¹⁶¹ M. Sharma,¹⁶¹ P. Sheldon,¹⁶¹ B. Snook,¹⁶¹ S. Tuo,¹⁶¹ J. Velkovska,¹⁶¹ M. W. Arenton,¹⁶² S. Boutle,¹⁶² B. Cox,¹⁶² B. Francis,¹⁶² J. Goodell,¹⁶² R. Hirosky,¹⁶² A. Ledovskoy,¹⁶² H. Li,¹⁶² C. Lin,¹⁶² C. Neu,¹⁶² J. Wood,¹⁶² S. Gollapinni,¹⁶³ R. Harr,¹⁶³ P. E. Karchin,¹⁶³ C. Kottachchi Kankanamge Don,¹⁶³ P. Lamichhane,¹⁶³ D. A. Belknap,¹⁶⁴ D. Carlsmith,¹⁶⁴ M. Cepeda,¹⁶⁴ S. Dasu,¹⁶⁴ S. Duric,¹⁶⁴ E. Friis,¹⁶⁴ R. Hall-Wilton,¹⁶⁴ M. Herndon,¹⁶⁴ A. Hervé,¹⁶⁴ P. Klabbers,¹⁶⁴ J. Klukas,¹⁶⁴ A. Lanaro,¹⁶⁴ C. Lazaridis,¹⁶⁴ A. Levine,¹⁶⁴ R. Loveless,¹⁶⁴ A. Mohapatra,¹⁶⁴ I. Ojalvo,¹⁶⁴ T. Perry,¹⁶⁴ G. A. Pierro,¹⁶⁴ G. Polese,¹⁶⁴ I. Ross,¹⁶⁴ T. Sarangi,¹⁶⁴ A. Savin,¹⁶⁴ W. H. Smith¹⁶⁴ and N. Woods¹⁶⁴

(CMS Collaboration)

- ¹*Yerevan Physics Institute, Yerevan, Armenia*
²*Institut für Hochenergiephysik der OeAW, Wien, Austria*
³*National Centre for Particle and High Energy Physics, Minsk, Belarus*
⁴*Universiteit Antwerpen, Antwerpen, Belgium*
⁵*Vrije Universiteit Brussel, Brussel, Belgium*
⁶*Université Libre de Bruxelles, Bruxelles, Belgium*
⁷*Ghent University, Ghent, Belgium*
⁸*Université Catholique de Louvain, Louvain-la-Neuve, Belgium*
⁹*Université de Mons, Mons, Belgium*
¹⁰*Centro Brasileiro de Pesquisas Físicas, Rio de Janeiro, Brazil*
¹¹*Universidade do Estado do Rio de Janeiro, Rio de Janeiro, Brazil*
^{12a}*Universidade Estadual Paulista, São Paulo, Brazil*
^{12b}*Universidade Federal do ABC, São Paulo, Brazil*
¹³*Institute for Nuclear Research and Nuclear Energy, Sofia, Bulgaria*
¹⁴*University of Sofia, Sofia, Bulgaria*
¹⁵*Institute of High Energy Physics, Beijing, China*
¹⁶*State Key Laboratory of Nuclear Physics and Technology, Peking University, Beijing, China*
¹⁷*Universidad de Los Andes, Bogota, Colombia*
¹⁸*Technical University of Split, Split, Croatia*
¹⁹*University of Split, Split, Croatia*
²⁰*Institute Rudjer Boskovic, Zagreb, Croatia*
²¹*University of Cyprus, Nicosia, Cyprus*
²²*Charles University, Prague, Czech Republic*
²³*Academy of Scientific Research and Technology of the Arab Republic of Egypt,
Egyptian Network of High Energy Physics, Cairo, Egypt*
²⁴*National Institute of Chemical Physics and Biophysics, Tallinn, Estonia*
²⁵*Department of Physics, University of Helsinki, Helsinki, Finland*
²⁶*Helsinki Institute of Physics, Helsinki, Finland*
²⁷*Lappeenranta University of Technology, Lappeenranta, Finland*
²⁸*DSM/IRFU, CEA/Saclay, Gif-sur-Yvette, France*
²⁹*Laboratoire Leprince-Ringuet, Ecole Polytechnique, IN2P3-CNRS, Palaiseau, France*
³⁰*Institut Pluridisciplinaire Hubert Curien, Université de Strasbourg,
Université de Haute Alsace Mulhouse, CNRS/IN2P3, Strasbourg, France*
³¹*Centre de Calcul de l'Institut National de Physique Nucleaire et de Physique des Particules,
CNRS/IN2P3, Villeurbanne, France*
³²*Université de Lyon, Université Claude Bernard Lyon 1, CNRS-IN2P3,
Institut de Physique Nucléaire de Lyon, Villeurbanne, France*
³³*Institute of High Energy Physics and Informatization, Tbilisi State University, Tbilisi, Georgia*
³⁴*RWTH Aachen University, I. Physikalisches Institut, Aachen, Germany*
³⁵*RWTH Aachen University, III. Physikalisches Institut A, Aachen, Germany*
³⁶*RWTH Aachen University, III. Physikalisches Institut B, Aachen, Germany*
³⁷*Deutsches Elektronen-Synchrotron, Hamburg, Germany*
³⁸*University of Hamburg, Hamburg, Germany*
³⁹*Institut für Experimentelle Kernphysik, Karlsruhe, Germany*
⁴⁰*Institute of Nuclear and Particle Physics (INPP), NCSR Demokritos, Aghia Paraskevi, Greece*
⁴¹*University of Athens, Athens, Greece*
⁴²*University of Ioánnina, Ioánnina, Greece*
⁴³*Wigner Research Centre for Physics, Budapest, Hungary*
⁴⁴*Institute of Nuclear Research ATOMKI, Debrecen, Hungary*
⁴⁵*University of Debrecen, Debrecen, Hungary*
⁴⁶*National Institute of Science Education and Research, Bhubaneswar, India*
⁴⁷*Panjab University, Chandigarh, India*
⁴⁸*University of Delhi, Delhi, India*
⁴⁹*Saha Institute of Nuclear Physics, Kolkata, India*
⁵⁰*Bhabha Atomic Research Centre, Mumbai, India*
⁵¹*Tata Institute of Fundamental Research, Mumbai, India*
⁵²*Institute for Research in Fundamental Sciences (IPM), Tehran, Iran*
⁵³*University College Dublin, Dublin, Ireland*
^{54a}*INFN Sezione di Bari, Bari, Italy*

- ^{54b}*Università di Bari, Bari, Italy*
^{54c}*Politecnico di Bari, Bari, Italy*
^{55a}*INFN Sezione di Bologna, Bologna, Italy*
^{55b}*Università di Bologna, Bologna, Italy*
^{56a}*INFN Sezione di Catania, Italy*
^{56b}*Università di Catania, Italy*
^{56c}*CSFNSM, Italy*
^{57a}*INFN Sezione di Firenze, Firenze, Italy*
^{57b}*Università di Firenze, Firenze, Italy*
⁵⁸*INFN Laboratori Nazionali di Frascati, Frascati, Italy*
^{59a}*INFN Sezione di Genova, Genova, Italy*
^{59b}*Università di Genova, Genova, Italy*
^{60a}*INFN Sezione di Milano-Bicocca, Milano, Italy*
^{60b}*Università di Milano-Bicocca, Milano, Italy*
^{61a}*INFN Sezione di Napoli, Napoli, Italy*
^{61b}*Università di Napoli 'Federico II', Napoli, Italy*
^{61c}*Università della Basilicata (Potenza), Napoli, Italy*
^{61d}*Università G. Marconi (Roma), Napoli, Italy*
^{62a}*INFN Sezione di Padova, Padova, Italy*
^{62b}*Università di Padova, Padova, Italy*
^{62c}*Università di Trento (Trento), Padova, Italy*
^{63a}*INFN Sezione di Pavia, Pavia, Italy*
^{63b}*Università di Pavia, Pavia, Italy*
^{64a}*INFN Sezione di Perugia, Perugia, Italy*
^{64b}*Università di Perugia, Perugia, Italy*
^{65a}*INFN Sezione di Pisa, Pisa, Italy*
^{65b}*Università di Pisa, Pisa, Italy*
^{65c}*Scuola Normale Superiore di Pisa, Pisa, Italy*
^{66a}*INFN Sezione di Roma, Roma, Italy*
^{66b}*Università di Roma, Roma, Italy*
^{67a}*INFN Sezione di Torino, Torino, Italy*
^{67b}*Università di Torino, Torino, Italy*
^{67c}*Università del Piemonte Orientale (Novara), Torino, Italy*
^{68a}*INFN Sezione di Trieste, Trieste, Italy*
^{68b}*Università di Trieste, Trieste, Italy*
⁶⁹*Kangwon National University, Chunchon, Korea*
⁷⁰*Kyungpook National University, Daegu, Korea*
⁷¹*Institute for Universe and Elementary Particles, Chonnam National University, Kwangju, Korea*
⁷²*Korea University, Seoul, Korea*
⁷³*University of Seoul, Seoul, Korea*
⁷⁴*Sungkyunkwan University, Suwon, Korea*
⁷⁵*Vilnius University, Vilnius, Lithuania*
⁷⁶*National Centre for Particle Physics, Universiti Malaya, Kuala Lumpur, Malaysia*
⁷⁷*Centro de Investigación y de Estudios Avanzados del IPN, Mexico City, Mexico*
⁷⁸*Universidad Iberoamericana, Mexico City, Mexico*
⁷⁹*Benemerita Universidad Autónoma de Puebla, Puebla, Mexico*
⁸⁰*Universidad Autónoma de San Luis Potosí, San Luis Potosí, Mexico*
⁸¹*University of Auckland, Auckland, New Zealand*
⁸²*University of Canterbury, Christchurch, New Zealand*
⁸³*National Centre for Physics, Quaid-I-Azam University, Islamabad, Pakistan*
⁸⁴*National Centre for Nuclear Research, Swierk, Poland*
⁸⁵*Institute of Experimental Physics, Faculty of Physics, University of Warsaw, Warsaw, Poland*
⁸⁶*Laboratório de Instrumentação e Física Experimental de Partículas, Lisboa, Portugal*
⁸⁷*Joint Institute for Nuclear Research, Dubna, Russia*
⁸⁸*Petersburg Nuclear Physics Institute, Gatchina (Saint Petersburg), Russia*
⁸⁹*Institute for Nuclear Research, Moscow, Russia*
⁹⁰*Institute for Theoretical and Experimental Physics, Moscow, Russia*
⁹¹*P.N. Lebedev Physical Institute, Moscow, Russia*
⁹²*Skobeltsyn Institute of Nuclear Physics, Lomonosov Moscow State University, Moscow, Russia*
⁹³*State Research Center of Russian Federation, Institute for High Energy Physics, Protvino, Russia*

- ⁹⁴*Faculty of Physics and Vinca Institute of Nuclear Sciences, University of Belgrade, Belgrade, Serbia*
- ⁹⁵*Centro de Investigaciones Energéticas Medioambientales y Tecnológicas (CIEMAT), Madrid, Spain*
- ⁹⁶*Universidad Autónoma de Madrid, Madrid, Spain*
- ⁹⁷*Universidad de Oviedo, Oviedo, Spain*
- ⁹⁸*Instituto de Física de Cantabria (IFCA), CSIC-Universidad de Cantabria, Santander, Spain*
- ⁹⁹*CERN, European Organization for Nuclear Research, Geneva, Switzerland*
- ¹⁰⁰*Paul Scherrer Institut, Villigen, Switzerland*
- ¹⁰¹*Institute for Particle Physics, ETH Zurich, Zurich, Switzerland*
- ¹⁰²*Universität Zürich, Zurich, Switzerland*
- ¹⁰³*National Central University, Chung-Li, Taiwan*
- ¹⁰⁴*National Taiwan University (NTU), Taipei, Taiwan*
- ¹⁰⁵*Chulalongkorn University, Bangkok, Thailand*
- ¹⁰⁶*Cukurova University, Adana, Turkey*
- ¹⁰⁷*Physics Department, Middle East Technical University, Ankara, Turkey*
- ¹⁰⁸*Bogazici University, Istanbul, Turkey*
- ¹⁰⁹*Istanbul Technical University, Istanbul, Turkey*
- ¹¹⁰*National Scientific Center, Kharkov Institute of Physics and Technology, Kharkov, Ukraine*
- ¹¹¹*University of Bristol, Bristol, United Kingdom*
- ¹¹²*Rutherford Appleton Laboratory, Didcot, United Kingdom*
- ¹¹³*Imperial College, London, United Kingdom*
- ¹¹⁴*Brunel University, Uxbridge, United Kingdom*
- ¹¹⁵*Baylor University, Waco, USA*
- ¹¹⁶*The University of Alabama, Tuscaloosa, USA*
- ¹¹⁷*Boston University, Boston, USA*
- ¹¹⁸*Brown University, Providence, USA*
- ¹¹⁹*University of California, Davis, Davis, USA*
- ¹²⁰*University of California, Los Angeles, USA*
- ¹²¹*University of California, Riverside, Riverside, USA*
- ¹²²*University of California, San Diego, La Jolla, USA*
- ¹²³*University of California, Santa Barbara, Santa Barbara, USA*
- ¹²⁴*California Institute of Technology, Pasadena, USA*
- ¹²⁵*Carnegie Mellon University, Pittsburgh, USA*
- ¹²⁶*University of Colorado at Boulder, Boulder, USA*
- ¹²⁷*Cornell University, Ithaca, USA*
- ¹²⁸*Fairfield University, Fairfield, USA*
- ¹²⁹*Fermi National Accelerator Laboratory, Batavia, USA*
- ¹³⁰*University of Florida, Gainesville, USA*
- ¹³¹*Florida International University, Miami, USA*
- ¹³²*Florida State University, Tallahassee, USA*
- ¹³³*Florida Institute of Technology, Melbourne, USA*
- ¹³⁴*University of Illinois at Chicago (UIC), Chicago, USA*
- ¹³⁵*The University of Iowa, Iowa City, USA*
- ¹³⁶*Johns Hopkins University, Baltimore, USA*
- ¹³⁷*The University of Kansas, Lawrence, USA*
- ¹³⁸*Kansas State University, Manhattan, USA*
- ¹³⁹*Lawrence Livermore National Laboratory, Livermore, USA*
- ¹⁴⁰*University of Maryland, College Park, USA*
- ¹⁴¹*Massachusetts Institute of Technology, Cambridge, USA*
- ¹⁴²*University of Minnesota, Minneapolis, USA*
- ¹⁴³*University of Mississippi, Oxford, USA*
- ¹⁴⁴*University of Nebraska-Lincoln, Lincoln, USA*
- ¹⁴⁵*State University of New York at Buffalo, Buffalo, USA*
- ¹⁴⁶*Northeastern University, Boston, USA*
- ¹⁴⁷*Northwestern University, Evanston, USA*
- ¹⁴⁸*University of Notre Dame, Notre Dame, USA*
- ¹⁴⁹*The Ohio State University, Columbus, USA*
- ¹⁵⁰*Princeton University, Princeton, USA*
- ¹⁵¹*University of Puerto Rico, Mayaguez, USA*
- ¹⁵²*Purdue University, West Lafayette, USA*
- ¹⁵³*Purdue University Calumet, Hammond, USA*

¹⁵⁴*Rice University, Houston, USA*¹⁵⁵*University of Rochester, Rochester, USA*¹⁵⁶*The Rockefeller University, New York, USA*¹⁵⁷*Rutgers, The State University of New Jersey, Piscataway, USA*¹⁵⁸*University of Tennessee, Knoxville, USA*¹⁵⁹*Texas A&M University, College Station, USA*¹⁶⁰*Texas Tech University, Lubbock, USA*¹⁶¹*Vanderbilt University, Nashville, USA*¹⁶²*University of Virginia, Charlottesville, USA*¹⁶³*Wayne State University, Detroit, USA*¹⁶⁴*University of Wisconsin, Madison, USA*^aDeceased.^bAlso at Vienna University of Technology, Vienna, Austria.^cAlso at CERN, European Organization for Nuclear Research, Geneva, Switzerland.^dAlso at Institut Pluridisciplinaire Hubert Curien, Université de Strasbourg, Université de Haute Alsace Mulhouse, CNRS/IN2P3, Strasbourg, France.^eAlso at National Institute of Chemical Physics and Biophysics, Tallinn, Estonia.^fAlso at Skobeltsyn Institute of Nuclear Physics, Lomonosov Moscow State University, Moscow, Russia.^gAlso at Universidade Estadual de Campinas, Campinas, Brazil.^hAlso at California Institute of Technology, Pasadena, USA.ⁱAlso at Laboratoire Leprince-Ringuet, Ecole Polytechnique, IN2P3-CNRS, Palaiseau, France.^jAlso at Suez University, Suez, Egypt.^kAlso at British University in Egypt, Cairo, Egypt.^lAlso at Fayoum University, El-Fayoum, Egypt.^mAlso at Ain Shams University, Cairo, Egypt.ⁿAlso at Université de Haute Alsace, Mulhouse, France.^oAlso at Joint Institute for Nuclear Research, Dubna, Russia.^pAlso at Brandenburg University of Technology, Cottbus, Germany.^qAlso at The University of Kansas, Lawrence, USA.^rAlso at Institute of Nuclear Research ATOMKI, Debrecen, Hungary.^sAlso at Eötvös Loránd University, Budapest, Hungary.^tAlso at University of Debrecen, Debrecen, Hungary.^uAlso at King Abdulaziz University, Jeddah, Saudi Arabia.^vAlso at University of Visva-Bharati, Santiniketan, India.^wAlso at University of Ruhuna, Matara, Sri Lanka.^xAlso at Isfahan University of Technology, Isfahan, Iran.^yAlso at Sharif University of Technology, Tehran, Iran.^zAlso at Plasma Physics Research Center, Science and Research Branch, Islamic Azad University, Tehran, Iran.^{aa}Also at Università degli Studi di Siena, Siena, Italy.^{bb}Also at Centre National de la Recherche Scientifique (CNRS)-IN2P3, Paris, France.^{cc}Also at Purdue University, West Lafayette, USA.^{dd}Also at Universidad Michoacana de San Nicolas de Hidalgo, Morelia, Mexico.^{ee}Also at Institute for Nuclear Research, Moscow, Russia.^{ff}Also at St. Petersburg State Polytechnical University, Saint Petersburg, Russia.^{gg}Also at Faculty of Physics, University of Belgrade, Belgrade, Serbia.^{hh}Also at Facoltà Ingegneria, Università di Roma, Roma, Italy.ⁱⁱAlso at Scuola Normale e Sezione dell'INFN, Pisa, Italy.^{jj}Also at University of Athens, Athens, Greece.^{kk}Also at Paul Scherrer Institut, Villigen, Switzerland.^{ll}Also at Institute for Theoretical and Experimental Physics, Moscow, Russia.^{mmm}Also at Albert Einstein Center for Fundamental Physics, Bern, Switzerland.ⁿⁿAlso at Gaziosmanpasa University, Tokat, Turkey.^{oo}Also at Adiyaman University, Adiyaman, Turkey.^{pp}Also at Cag University, Mersin, Turkey.^{qq}Also at Mersin University, Mersin, Turkey.^{rr}Also at Izmir Institute of Technology, Izmir, Turkey.^{ss}Also at Ozyegin University, Istanbul, Turkey.^{tt}Also at Kafkas University, Kars, Turkey.^{uu}Also at Mimar Sinan University, Istanbul, Turkey.

^{vv} Also at Rutherford Appleton Laboratory, Didcot, United Kingdom.

^{ww} Also at School of Physics and Astronomy, University of Southampton, Southampton, United Kingdom.

^{xx} Also at University of Belgrade, Faculty of Physics and Vinca Institute of Nuclear Sciences, Belgrade, Serbia.

^{yy} Also at Argonne National Laboratory, Argonne, USA.

^{zz} Also at Erzincan University, Erzincan, Turkey.

^{aaa} Also at Yildiz Technical University, Istanbul, Turkey.

^{bbb} Also at Texas A&M University at Qatar, Doha, Qatar.

^{ccc} Also at Kyungpook National University, Daegu, Korea.